## Appendices

## A. Physical Constants

We use SI units throughout this text. Simple ways to convert between SI and other popular units, such as Gaussian, may be found in Refs. [105-108].

The Committee on Data for Science and Technology (CODATA) of NIST maintains the values of many physical constants [92]. The most current values can be obtained from the CODATA web site [820]. Some commonly used constants are listed below:

| quantity | symbol | value | units |
| :--- | :---: | :--- | :--- |
| speed of light in vacuum | $c_{0}, c$ | 299792458 | $\mathrm{~m} \mathrm{~s}^{-1}$ |
| permittivity of vacuum | $\epsilon_{0}$ | $8.854187817 \times 10^{-12}$ | $\mathrm{~F} \mathrm{~m}^{-1}$ |
| permeability of vacuum | $\mu_{0}$ | $4 \pi \times 10^{-7}$ | $\mathrm{H} \mathrm{m}^{-1}$ |
| characteristic impedance | $\eta_{0}, Z_{0}$ | 376.730313461 | $\Omega$ |
| electron charge | $e$ | $1.602176462 \times 10^{-19}$ | C |
| electron mass | $m_{e}$ | $9.109381887 \times 10^{-31}$ | kg |
| Boltzmann constant | $k$ | $1.380650324 \times 10^{-23}$ | $\mathrm{~J} \mathrm{~K}^{-1}$ |
| Avogadro constant | $N_{A}, L$ | $6.022141994 \times 10^{23}$ | $\mathrm{~mol}^{-1}$ |
| Planck constant | $h$ | $6.62606876 \times 10^{-34}$ | $\mathrm{~J} / \mathrm{Hz}^{2}$ |
|  | $G$ | $6.67259 \times 10^{-11}$ | $\mathrm{~m}^{3} \mathrm{~kg}^{-1} \mathrm{~s}^{-2}$ |
| Gravitational constant | $G$ | $5.972 \times 10^{24}$ | $\mathrm{~kg}^{2}$ |
| Earth mass | $M_{\oplus}$ | $a_{e}$ | 6378 |
| Earth equatorial radius |  |  | $\mathrm{km}^{2}$ |

In the table, the constants $c, \mu_{0}$ are taken to be exact, whereas $\epsilon_{0}, \eta_{0}$ are derived from the relationships:

$$
\epsilon_{0}=\frac{1}{\mu_{0} c^{2}}, \quad \eta_{0}=\sqrt{\frac{\mu_{0}}{\epsilon_{0}}}=\mu_{0} c
$$

The energy unit of electron volt $(\mathrm{eV})$ is defined to be the work done by an electron in moving across a voltage of one volt, that is, $1 \mathrm{eV}=1.602176462 \times 10^{-19} \mathrm{C} \cdot 1 \mathrm{~V}$, or

$$
1 \mathrm{eV}=1.602176462 \times 10^{-19} \mathrm{~J}
$$

In units of $\mathrm{eV} / \mathrm{Hz}$, Planck's constant $h$ is:

$$
h=4.13566727 \times 10^{-15} \mathrm{eV} / \mathrm{Hz}=1 \mathrm{eV} / 241.8 \mathrm{THz}
$$

that is, 1 eV corresponds to a frequency of 241.8 THz , or a wavelength of $1.24 \mu \mathrm{~m}$.

## B. Electromagnetic Frequency Bands

The $\mathrm{ITU}^{\dagger}$ divides the radio frequency (RF) spectrum into the following frequency and wavelength bands in the range from 30 Hz to 3000 GHz :

| RF Spectrum |  |  |  |  |  |
| :--- | :--- | :---: | :--- | :---: | :--- |
|  | band designations | frequency | wavelength |  |  |
| ELF | Extremely Low Frequency | $30-300$ | Hz | $1-10$ | Mm |
| VF | Voice Frequency | $300-3000$ | Hz | $100-1000$ | km |
| VLF | Very Low Frequency | $3-30$ | kHz | $10-100$ | km |
| LF | Low Frequency | $30-300$ | kHz | $1-10$ | km |
| MF | Medium Frequency | $300-3000$ | kHz | $100-1000$ | m |
| HF | High Frequency | $3-30$ | MHz | $10-100$ | m |
| VHF | Very High Frequency | $30-300$ | MHz | $1-10$ | m |
| UHF | Ultra High Frequency | $300-3000$ | MHz | $10-100$ | cm |
| SHF | Super High Frequency | $3-30$ | GHz | $1-10$ | cm |
| EHF | Extremely High Frequency | $30-300$ | GHz | $1-10$ | mm |
|  | Submillimeter | $300-3000$ | GHz | $100-1000$ | $\mu \mathrm{~m}$ |

An alternative subdivision of the low-frequency bands is to designate the bands $3-30 \mathrm{~Hz}, 30-300 \mathrm{~Hz}$, and $300-3000 \mathrm{~Hz}$ as extremely low frequency (ELF), super low frequency (SLF), and ultra low frequency (ULF), respectively.

Microwaves span the $300 \mathrm{MHz}-300 \mathrm{GHz}$ frequency range. Typical microwave and satellite communication systems and radar use the $1-30 \mathrm{GHz}$ band. The $30-300 \mathrm{GHz}$ EHF band is also referred to as the millimeter band.

The $1-100 \mathrm{GHz}$ range is subdivided further into the subbands shown on the right.

| Microwave Bands |  |  |
| :---: | :---: | :---: |
| band | frequ | ency |
| L | 1-2 | GHz |
| S | 2-4 | GHz |
| C | 4-8 | GHz |
| X | 8-12 | GHz |
| Ku | 12-18 | GHz |
| K | 18-27 | GHz |
| Ka | 27-40 | GHz |
| V | 40-75 | GHz |
| W | 80-100 | GHz |

Some typical RF applications are as follows. AM radio is broadcast at 535-1700 kHz falling within the MF band. The HF band is used in short-wave radio, navigation, amateur, and CB bands. FM radio at $88-108 \mathrm{MHz}$, ordinary TV, police, walkie-talkies, and remote control occupy the VHF band.

Cell phones, personal communication systems (PCS), pagers, cordless phones, global positioning systems (GPS), RF identification systems (RFID), UHF-TV channels, microwave ovens, and long-range surveillance radar fall within the UHF band.

[^0]The SHF microwave band is used in radar (traffic control, surveillance, tracking, missile guidance, mapping, weather), satellite communications, direct-broadcast satellite (DBS), and microwave relay systems. Multipoint multichannel (MMDS) and local multipoint (LMDS) distribution services, fall within UHF and SHF at 2.5 GHz and 30 GHz .

Industrial, scientific, and medical (ISM) bands are within the UHF and low SHF, at 900 $\mathrm{MHz}, 2.4 \mathrm{GHz}$, and 5.8 GHz . Radio astronomy occupies several bands, from UHF to L-W microwave bands.

Beyond RF, come the infrared (IR), visible, ultraviolet (UV), X-ray, and $\gamma$-ray bands. The IR range extends over $3-300 \mathrm{THz}$, or $1-100 \mu \mathrm{~m}$. Many IR applications fall in the $1-20 \mu \mathrm{~m}$ band. For example, optical fiber communications typically use laser light at $1.55 \mu \mathrm{~m}$ or 193 THz because of the low fiber losses at that frequency. The UV range lies beyond the visible band, extending typically over $10-400 \mathrm{~nm}$.

| band | wavelength | frequency | energy |
| :--- | :---: | :---: | :---: |
| infrared | $100-1 \mu \mathrm{~m}$ | $3-300 \mathrm{THz}$ |  |
| ultraviolet | $400-10 \mathrm{~nm}$ | $750 \mathrm{THz}-30 \mathrm{PHz}$ |  |
| X-Ray | $10 \mathrm{~nm}-100 \mathrm{pm}$ | $30 \mathrm{PHz}-3 \mathrm{EHz}$ | $0.124-124 \mathrm{keV}$ |
| $\gamma$-ray | $<100 \mathrm{pm}$ | $>3 \mathrm{EHz}$ | $>124 \mathrm{keV}$ |

The $\mathrm{CIE}^{\dagger}$ defines the visible spectrum to be the wavelength range $380-780 \mathrm{~nm}$, or 385-789 THz. Colors fall within the following typical wavelength/frequency ranges:

|  | Visible Spectrum |  |
| :--- | :--- | :---: |
| color | wavelength | frequency |
| red | $780-620 \mathrm{~nm}$ | $385-484 \mathrm{THz}$ |
| orange | $620-600 \mathrm{~nm}$ | $484-500 \mathrm{THz}$ |
| yellow | $600-580 \mathrm{~nm}$ | $500-517 \mathrm{THz}$ |
| green | $580-490 \mathrm{~nm}$ | $517-612 \mathrm{THz}$ |
| blue | $490-450 \mathrm{~nm}$ | $612-667 \mathrm{THz}$ |
| violet | $450-380 \mathrm{~nm}$ | $667-789 \mathrm{THz}$ |

X -ray frequencies fall in the PHz (petahertz) range and $\gamma$-ray frequencies in the EHz (exahertz) range. ${ }^{\ddagger} \mathrm{X}$-rays and $\gamma$-rays are best described in terms of their energy, which is related to frequency through Planck's relationship, $E=h f$. X-rays have typical energies of the order of keV , and $\gamma$-rays, of the order of MeV and beyond. By comparison, photons in the visible spectrum have energies of a couple of eV .

The earth's atmosphere is mostly opaque to electromagnetic radiation, except for three significant "windows", the visible, the infrared, and the radio windows. These three bands span the wavelength ranges of $380-780 \mathrm{~nm}, 1-12 \mu \mathrm{~m}$, and $5 \mathrm{~mm}-20 \mathrm{~m}$, respectively.

Within the 1-10 $\mu \mathrm{m}$ infrared band there are some narrow transparent windows. For the rest of the IR range ( $1-1000 \mu \mathrm{~m}$ ), water and carbon dioxide molecules absorb infrared radiation-this is responsible for the Greenhouse effect. There are also some minor transparent windows for 17-40 and 330-370 $\mu \mathrm{m}$.

[^1]Beyond the visible band, ultraviolet and X-ray radiation are absorbed by ozone and molecular oxygen (except for the ozone holes.)

## C. Vector Identities and Integral Theorems

## Algebraic Identities

$$
\begin{align*}
|\boldsymbol{A}|^{2}|\boldsymbol{B}|^{2} & =|\boldsymbol{A} \cdot \boldsymbol{B}|^{2}+|\boldsymbol{A} \times \boldsymbol{B}|^{2}  \tag{C.1}\\
(\boldsymbol{A} \times \boldsymbol{B}) \cdot \boldsymbol{C} & =(\boldsymbol{B} \times \boldsymbol{C}) \cdot \boldsymbol{A}=(\boldsymbol{C} \times \boldsymbol{A}) \cdot \boldsymbol{B}  \tag{C.2}\\
\boldsymbol{A} \times(\boldsymbol{B} \times \boldsymbol{C}) & =\boldsymbol{B}(\boldsymbol{A} \cdot \boldsymbol{C})-\boldsymbol{C}(\boldsymbol{A} \cdot \boldsymbol{B})  \tag{C.3}\\
(\boldsymbol{A} \times \boldsymbol{B}) \cdot(\boldsymbol{C} \times \boldsymbol{D}) & =(\boldsymbol{A} \cdot \boldsymbol{C})(\boldsymbol{B} \cdot \boldsymbol{D})-(\boldsymbol{A} \cdot \boldsymbol{D})(\boldsymbol{B} \cdot \boldsymbol{C})  \tag{.4}\\
(\boldsymbol{A} \times \boldsymbol{B}) \times(\boldsymbol{C} \times \boldsymbol{D}) & =[(\boldsymbol{A} \times \boldsymbol{B}) \cdot \boldsymbol{D}] \boldsymbol{C}-[(\boldsymbol{A} \times \boldsymbol{B}) \cdot \boldsymbol{C}] \boldsymbol{D}  \tag{C.5}\\
\boldsymbol{A} & =\hat{\mathbf{n}} \times(\boldsymbol{A} \times \hat{\mathbf{n}})+(\hat{\mathbf{n}} \cdot \boldsymbol{A}) \hat{\mathbf{n}}=\boldsymbol{A}_{\perp}+\boldsymbol{A}_{\|} \tag{C.6}
\end{align*}
$$

where $\hat{\mathbf{n}}$ is any unit vector, and $\boldsymbol{A}_{\perp}, \boldsymbol{A}_{\|}$are the components of $\boldsymbol{A}$ perpendicular and parallel to $\hat{\mathbf{n}}$. Note also that $\hat{\mathbf{n}} \times(\boldsymbol{A} \times \hat{\mathbf{n}})=(\hat{\mathbf{n}} \times \boldsymbol{A}) \times \hat{\mathbf{n}}$. A three-dimensional vector can equally well be represented as a column vector:

$$
\boldsymbol{a}=a_{x} \hat{\mathbf{x}}+a_{y} \hat{\mathbf{y}}+a_{z} \hat{\mathbf{z}} \quad \Leftrightarrow \quad \boldsymbol{a}=\left[\begin{array}{l}
a_{x}  \tag{C.7}\\
a_{y} \\
b_{z}
\end{array}\right]
$$

Consequently, the dot and cross products may be represented in matrix form:

$$
\begin{array}{cc}
\boldsymbol{a} \cdot \boldsymbol{b} \quad \Leftrightarrow \quad \boldsymbol{a}^{T} \boldsymbol{b}=\left[a_{x}, a_{y}, a_{z}\right]\left[\begin{array}{l}
b_{x} \\
b_{y} \\
b_{z}
\end{array}\right]=a_{x} b_{x}+a_{y} b_{y}+a_{z} b_{z} \\
\boldsymbol{a} \times \boldsymbol{b} & \Leftrightarrow \quad A \boldsymbol{b}=\left[\begin{array}{rrr}
0 & -a_{z} & a_{y} \\
a_{z} & 0 & -a_{x} \\
-a_{y} & a_{x} & 0
\end{array}\right]\left[\begin{array}{l}
b_{x} \\
b_{y} \\
b_{z}
\end{array}\right]=\left[\begin{array}{l}
a_{y} b_{z}-a_{z} b_{y} \\
a_{z} b_{x}-a_{x} b_{z} \\
a_{x} b_{y}-a_{y} b_{x}
\end{array}\right] \tag{С.9}
\end{array}
$$

The cross-product matrix $A$ satisfies the following identity:

$$
\begin{equation*}
A^{2}=\boldsymbol{a} \boldsymbol{a}^{T}-\left(\boldsymbol{a}^{T} \boldsymbol{a}\right) I \tag{С.10}
\end{equation*}
$$

where $I$ is the $3 \times 3$ identity matrix. Applied to a unit vector $\hat{\mathbf{n}}$, this identity reads:

$$
I=\hat{\mathbf{n}} \hat{\mathbf{n}}^{T}-\hat{N}^{2}, \quad \text { where } \quad \hat{\mathbf{n}}=\left[\begin{array}{c}
\hat{n}_{x}  \tag{C.11}\\
\hat{n}_{y} \\
\hat{n}_{z}
\end{array}\right], \quad \hat{N}=\left[\begin{array}{ccc}
0 & -\hat{n}_{z} & \hat{n}_{y} \\
\hat{n}_{z} & 0 & -\hat{n}_{X} \\
-\hat{n}_{y} & \hat{n}_{x} & 0
\end{array}\right], \quad \hat{\mathbf{n}}^{T} \hat{\mathbf{n}}=1
$$

This corresponds to the matrix form of the parallel/transverse decomposition (C.6). Indeed, we have $\boldsymbol{a}_{\|}=\hat{\mathbf{n}}\left(\hat{\mathbf{n}}^{T} \boldsymbol{a}\right)$ and $\boldsymbol{a}_{\perp}=(\hat{\mathbf{n}} \times \boldsymbol{a}) \times \hat{\mathbf{n}}=-\hat{\mathbf{n}} \times(\hat{\mathbf{n}} \times \boldsymbol{a})=-\hat{N}(\hat{N} \boldsymbol{a})=-\hat{N}^{2} \boldsymbol{a}$. Therefore, $\boldsymbol{a}=I \boldsymbol{a}=\left(\hat{\mathbf{n}} \hat{\mathbf{n}}^{T}-\hat{N}^{2}\right) \boldsymbol{a}=\boldsymbol{a}_{\|}+\boldsymbol{a}_{\perp}$.

## Differential Identities

$$
\begin{align*}
\nabla \times(\nabla \psi) & =0  \tag{C.12}\\
\nabla \cdot(\nabla \times \boldsymbol{A}) & =0  \tag{С.13}\\
\nabla \cdot(\psi \boldsymbol{A}) & =\boldsymbol{A} \cdot \nabla \psi+\psi \nabla \cdot \boldsymbol{A}  \tag{С.14}\\
\nabla \times(\boldsymbol{A}) & =\psi \nabla \times \boldsymbol{A}+\nabla \psi \times \boldsymbol{A}  \tag{С.15}\\
\nabla(\boldsymbol{A} \cdot \boldsymbol{B}) & =(\boldsymbol{A} \cdot \nabla) \boldsymbol{B}+(\boldsymbol{B} \cdot \nabla) \boldsymbol{A}+\boldsymbol{A} \times(\nabla \times \boldsymbol{B})+\boldsymbol{B} \times(\nabla \times \boldsymbol{A})  \tag{C.16}\\
\nabla \cdot(\boldsymbol{A} \times \boldsymbol{B}) & =\boldsymbol{B} \cdot(\nabla \times \boldsymbol{A})-\boldsymbol{A} \cdot(\nabla \times \boldsymbol{B})  \tag{C.17}\\
\nabla \times(\boldsymbol{A} \times \boldsymbol{B}) & =\boldsymbol{A}(\nabla \cdot \boldsymbol{B})-\boldsymbol{B}(\nabla \cdot \boldsymbol{A})+(\boldsymbol{B} \cdot \nabla) \boldsymbol{A}-(\boldsymbol{A} \cdot \nabla) \boldsymbol{B}  \tag{С.18}\\
\nabla \times(\nabla \times \boldsymbol{A}) & =\nabla(\nabla \cdot \boldsymbol{A})-\nabla^{2} \boldsymbol{A} \tag{С.19}
\end{align*}
$$

$$
\begin{aligned}
& A_{x} \nabla B_{x}+A_{y} \nabla B_{y}+A_{z} \nabla B_{z}=(\boldsymbol{A} \cdot \nabla) \boldsymbol{B}+\boldsymbol{A} \times(\nabla \times \boldsymbol{B}) \\
& B_{x} \nabla A_{x}+B_{y} \nabla A_{y}+B_{z} \nabla A_{z}=(\boldsymbol{B} \cdot \nabla) \boldsymbol{A}+\boldsymbol{B} \times(\boldsymbol{\nabla} \times \boldsymbol{A})
\end{aligned}
$$

$$
(\hat{\mathbf{n}} \times \nabla) \times A=\hat{\mathbf{n}} \times(\nabla \times A)+(\hat{\mathbf{n}} \cdot \nabla) A-\hat{\mathbf{n}}(\nabla \cdot A)
$$

$$
\begin{align*}
\psi(\hat{\mathbf{n}} \cdot \nabla) & \boldsymbol{E}-\boldsymbol{E}(\hat{\mathbf{n}} \cdot \nabla \psi)=[(\hat{\mathbf{n}} \cdot \nabla)(\psi \boldsymbol{E})+\hat{\mathbf{n}} \times(\nabla \times(\psi \boldsymbol{E}))-\hat{\mathbf{n}} \nabla \cdot(\psi \boldsymbol{E})] \\
+ & {[\hat{\mathbf{n}} \psi \nabla \cdot \boldsymbol{E}-(\hat{\mathbf{n}} \times \boldsymbol{E}) \times \nabla \psi-\psi \hat{\mathbf{n}} \times(\nabla \times \boldsymbol{E})-(\hat{\mathbf{n}} \cdot \boldsymbol{E}) \nabla \psi] } \tag{С.23}
\end{align*}
$$

With $\mathbf{r}=x \hat{\mathbf{x}}+y \hat{\mathbf{y}}+z \hat{\mathbf{z}}, r=|\mathbf{r}|=\sqrt{x^{2}+y^{2}+z^{2}}$, and the unit vector $\hat{\mathbf{r}}=\mathbf{r} / r$, we have:

$$
\begin{equation*}
\nabla r=\hat{\mathbf{r}}, \quad \nabla r^{2}=2 \mathbf{r}, \quad \nabla \frac{1}{r}=-\frac{\hat{\mathbf{r}}}{r^{2}}, \quad \nabla \cdot \mathbf{r}=3, \quad \nabla \times \mathbf{r}=0, \quad \nabla \cdot \hat{\mathbf{r}}=\frac{2}{r} \tag{С.24}
\end{equation*}
$$

## Integral Theorems for Closed Surfaces

The theorems involve a volume $V$ surrounded by a closed surface $S$. The divergence or Gauss' theorem is:

$$
\begin{equation*}
\int_{V} \boldsymbol{\nabla} \cdot \boldsymbol{A} d V=\oint_{S} \boldsymbol{A} \cdot \hat{\mathbf{n}} d S \quad \text { (Gauss' divergence theorem) } \tag{C.25}
\end{equation*}
$$

where $\hat{\mathbf{n}}$ is the outward normal to the surface. Green's first and second identities are:

$$
\begin{gather*}
\int_{V}\left[\varphi \nabla^{2} \psi+\nabla \varphi \cdot \nabla \psi\right] d V=\oint_{S} \varphi \frac{\partial \psi}{\partial n} d S  \tag{C.26}\\
\int_{V}\left[\varphi \nabla^{2} \psi-\psi \nabla^{2} \varphi\right] d V=\oint_{S}\left(\varphi \frac{\partial \psi}{\partial n}-\psi \frac{\partial \varphi}{\partial n}\right) d S \tag{C.27}
\end{gather*}
$$

where $\frac{\partial}{\partial n}=\hat{\mathbf{n}} \cdot \nabla$ is the directional derivative along $\hat{\mathbf{n}}$. Some related theorems are:

$$
\begin{gather*}
\int_{V} \nabla^{2} \psi d V=\oint_{S} \hat{\mathbf{n}} \cdot \nabla \psi d S=\oint_{S} \frac{\partial \psi}{\partial n} d S  \tag{C.28}\\
\int_{V} \nabla \psi d V=\oint_{S} \psi \hat{\mathbf{n}} d S  \tag{C.29}\\
\int_{V} \nabla^{2} \boldsymbol{A} d V=\oint_{S}(\hat{\mathbf{n}} \cdot \nabla) \boldsymbol{A} d S=\oint_{S} \frac{\partial \boldsymbol{A}}{\partial n} d S  \tag{С.30}\\
\oint_{S}(\hat{\mathbf{n}} \times \nabla) \times \boldsymbol{A} d S=\oint_{S}[\hat{\mathbf{n}} \times(\nabla \times \boldsymbol{A})+(\hat{\mathbf{n}} \cdot \nabla) \boldsymbol{A}-\hat{\mathbf{n}}(\nabla \cdot \boldsymbol{A})] d S=0  \tag{C.31}\\
\text { Using Eqs. (C.23) and (C.31), we find: }  \tag{С.32}\\
\int_{V} \nabla \times \boldsymbol{A} d V=\oint_{S} \hat{\mathbf{n}} \times \boldsymbol{A} d S \\
\oint_{S}\left(\psi \frac{\partial \boldsymbol{E}}{\partial n}-\boldsymbol{E} \frac{\partial \psi}{\partial n}\right) d S=  \tag{C.33}\\
=\oint_{S}[\hat{\mathbf{n}} \psi \nabla \cdot \boldsymbol{E}-(\hat{\mathbf{n}} \times \boldsymbol{E}) \times \nabla \psi-\psi \hat{\mathbf{n}} \times(\nabla \times \boldsymbol{E})-(\hat{\mathbf{n}} \cdot \boldsymbol{E}) \nabla \psi] d S
\end{gather*}
$$

The vectorial forms of Green's identities are [708,705]:

$$
\begin{equation*}
\int_{V}(\boldsymbol{\nabla} \times \boldsymbol{A} \cdot \boldsymbol{\nabla} \times \boldsymbol{B}-\boldsymbol{A} \cdot \boldsymbol{\nabla} \times \boldsymbol{\nabla} \times \boldsymbol{B}) d V=\oint_{S} \hat{\mathbf{n}} \cdot(\boldsymbol{A} \times \nabla \times \boldsymbol{B}) d S \tag{С.34}
\end{equation*}
$$

$$
\begin{equation*}
\int_{V}(\boldsymbol{B} \cdot \nabla \times \nabla \times \boldsymbol{A}-\boldsymbol{A} \cdot \nabla \times \boldsymbol{\nabla} \times \boldsymbol{B}) d V=\oint_{S} \hat{\mathbf{n}} \cdot(\boldsymbol{A} \times \nabla \times \boldsymbol{B}-\boldsymbol{B} \times \nabla \times \boldsymbol{A}) d S \tag{C.35}
\end{equation*}
$$

## Integral Theorems for Open Surfaces

Stokes' theorem involves an open surface $S$ and its boundary contour $C$ :

$$
\begin{equation*}
\int_{S} \hat{\mathbf{n}} \cdot \boldsymbol{\nabla} \times \boldsymbol{A} d S=\oint_{C} \boldsymbol{A} \cdot d \mathbf{l} \quad \text { (Stokes' theorem) } \tag{С.36}
\end{equation*}
$$

where $d \mathbf{l}$ is the tangential path length around $C$. Some related theorems are:

$$
\begin{gather*}
\int_{S}[\psi \hat{\mathbf{n}} \cdot \boldsymbol{\nabla} \times \boldsymbol{A}-(\hat{\mathbf{n}} \times \boldsymbol{A}) \cdot \nabla \psi] d S=\oint_{C} \psi \boldsymbol{A} \cdot d \mathbf{l}  \tag{C.37}\\
\int_{S}[(\boldsymbol{\nabla} \psi) \hat{\mathbf{n}} \cdot \nabla \times \boldsymbol{A}-((\hat{\mathbf{n}} \times \boldsymbol{A}) \cdot \nabla) \nabla \psi] d S=\oint_{C}(\nabla \psi) \boldsymbol{A} \cdot d \mathbf{l}  \tag{C.38}\\
\int_{S} \hat{\mathbf{n}} \times \nabla \psi d S=\oint_{C} \psi d \mathbf{l} \tag{С.39}
\end{gather*}
$$

$$
\begin{gather*}
\int_{S}(\hat{\mathbf{n}} \times \nabla) \times \boldsymbol{A} d S=\int_{S}[\hat{\mathbf{n}} \times(\nabla \times \boldsymbol{A})+(\hat{\mathbf{n}} \cdot \nabla) \boldsymbol{A}-\hat{\mathbf{n}}(\boldsymbol{\nabla} \cdot \boldsymbol{A})] d S=\oint_{C} d \mathbf{l} \times \boldsymbol{A} \\
\int_{S} \hat{\mathbf{n}} d S=\frac{1}{2} \oint_{C} \mathbf{r} \times d \mathbf{l} \tag{C.41}
\end{gather*}
$$

Eq. (C.41) is a special case of (C.40). Using Eqs. (C.23) and (C.40) we find:

$$
\begin{align*}
\int_{S}(\psi & \left.\frac{\partial \boldsymbol{E}}{\partial n}-\boldsymbol{E} \frac{\partial \psi}{\partial n}\right) d S+\oint_{C} \psi \boldsymbol{E} \times d \mathbf{l}= \\
& =\int_{S}[\hat{\mathbf{n}} \psi \nabla \cdot \boldsymbol{E}-(\hat{\mathbf{n}} \times \boldsymbol{E}) \times \nabla \psi-\psi \hat{\mathbf{n}} \times(\nabla \times \boldsymbol{E})-(\hat{\mathbf{n}} \cdot \boldsymbol{E}) \nabla \psi] d S \tag{C.42}
\end{align*}
$$

## D. Green's Functions

The Green's functions for the Laplace, Helmholtz, and one-dimensional Helmholtz equations are listed below:

$$
\begin{gather*}
\nabla^{2} g(\mathbf{r})=-\delta^{(3)}(\mathbf{r}) \Rightarrow g(\mathbf{r})=\frac{1}{4 \pi r}  \tag{D.1}\\
\left(\nabla^{2}+k^{2}\right) G(\mathbf{r})=-\delta^{(3)}(\mathbf{r}) \Rightarrow G(\mathbf{r})=\frac{e^{-j k r}}{4 \pi r}  \tag{D.2}\\
\left(\partial_{z}^{2}+\beta^{2}\right) g(z)=-\delta(z) \Rightarrow g(z)=\frac{e^{-j \beta|z|}}{2 j \beta} \tag{D.3}
\end{gather*}
$$

where $r=|\mathbf{r}|$. Eqs. (D.2) and (D.3) are appropriate for describing outgoing waves. We considered other versions of (D.3) in Sec. 20.3. A more general identity satisfied by the Green's function $g(\mathbf{r})$ of Eq. (D.1) is as follows (for a proof, see Refs. [114,115]):

$$
\begin{equation*}
\partial_{i} \partial_{j} g(\mathbf{r})=-\frac{1}{3} \delta_{i j} \delta^{(3)}(\mathbf{r})+\frac{3 x_{i} x_{j}-r^{2} \delta_{i j}}{r^{4}} g(\mathbf{r}) \quad i, j=1,2,3 \tag{D.4}
\end{equation*}
$$

where $\partial_{i}=\partial / \partial x_{i}$ and $x_{i}$ stands for any of $x, y, z$. By summing the $i, j$ indices, Eq. (D.4) reduces to (D.1). Using this identity, we find for the Green's function $G(\mathbf{r})=e^{-j k r} / 4 \pi r$ :

$$
\begin{equation*}
\partial_{i} \partial_{j} G(\mathbf{r})=-\frac{1}{3} \delta_{i j} \delta^{(3)}(\mathbf{r})+\left[\left(j k+\frac{1}{r}\right) \frac{3 x_{i} x_{j}-r^{2} \delta_{i j}}{r^{3}}-k^{2} \frac{x_{i} x_{j}}{r^{2}}\right] G(\mathbf{r}) \tag{D.5}
\end{equation*}
$$

This reduces to Eq. (D.2) upon summing the indices. For any fixed vector p, Eq. (D.5) is equivalent to the vectorial identity:

$$
\boldsymbol{\nabla} \times \boldsymbol{\nabla} \times[\mathbf{p} G(\mathbf{r})]=\frac{2}{3} \mathbf{p} \delta^{(3)}(\mathbf{r})+\left[\left(j k+\frac{1}{r}\right) \frac{3 \hat{\mathbf{r}}(\hat{\mathbf{r}} \cdot \mathbf{p})-\mathbf{p}}{r^{2}}+k^{2} \hat{\mathbf{r}} \times(\mathbf{p} \times \hat{\mathbf{r}})\right] G(\mathbf{r})
$$

The second term on the right is simply the left-hand side evaluated at points away from the origin, thus, we may write:

$$
\begin{equation*}
\nabla \times \nabla \times[\mathbf{p} G(\mathbf{r})]=\frac{2}{3} \mathbf{p} \delta^{(3)}(\mathbf{r})+[\nabla \times \nabla \times[\mathbf{p} G(\mathbf{r})]]_{\mathbf{r} \neq 0} \tag{D.7}
\end{equation*}
$$

Then, Eq. (D.7) implies the following integrated identity, where $\boldsymbol{\nabla}$ is with respect to $\mathbf{r}$ :
$\boldsymbol{\nabla} \times \nabla \times \int_{V} \boldsymbol{P}\left(\mathbf{r}^{\prime}\right) G\left(\mathbf{r}-\mathbf{r}^{\prime}\right) d V^{\prime}=\frac{2}{3} \boldsymbol{P}(\mathbf{r})+\int_{V}\left[\nabla \times \boldsymbol{\nabla} \times\left[\boldsymbol{P}\left(\mathbf{r}^{\prime}\right) G\left(\mathbf{r}-\mathbf{r}^{\prime}\right)\right]\right]_{\mathbf{r}^{\prime} \neq \mathbf{r}} d V^{\prime}$ (D.8) and $\mathbf{r}$ is assumed to lie within $V$. If $\mathbf{r}$ is outside $V$, then the term $2 \boldsymbol{P}(\mathbf{r}) / 3$ is absent.

Technically, the integrals in (D.8) are principal-value integrals, that is, the limits as $\delta \rightarrow 0$ of the integrals over $V-V_{\delta}(\mathbf{r})$, where $V_{\delta}(\mathbf{r})$ is an excluded small sphere of radius $\delta$ centered about $\mathbf{r}$. The $2 \boldsymbol{P}(\mathbf{r}) / 3$ term has a different form if the excluded volume $V_{\delta}(\mathbf{r})$ has shape other than a sphere or a cube. See Refs. [27,143,155,205] and [109-113] for the definitions and properties of such principal value integrals.

Another useful result is the so-called Weyl representation or plane-wave-spectrum representation $[22,26-28,198]$ of the outgoing Helmholtz Green's function $G(\mathbf{r})$ :

$$
\begin{equation*}
G(\mathbf{r})=\frac{e^{-j k r}}{4 \pi r}=\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \frac{e^{-j\left(k_{x} x+k_{y} y\right)} e^{-j k_{z}|z|}}{2 j k_{z}} \frac{d k_{x} d k_{y}}{(2 \pi)^{2}} \tag{D.9}
\end{equation*}
$$

where $k_{z}^{2}=k^{2}-k_{\perp}^{2}$, with $k_{\perp}=\sqrt{k_{x}^{2}+k_{y}^{2}}$. In order to correspond to either outgoing waves or decaying evanescent waves, $k_{Z}$ must be defined more precisely as follows:

$$
k_{z}=\left\{\begin{array}{rll}
\sqrt{k^{2}-k_{\perp}^{2}}, & \text { if } \quad k_{\perp} \leq k, & \text { (propagating modes) }  \tag{D.10}\\
-j \sqrt{k_{\perp}^{2}-k^{2}}, & \text { if } \quad k_{\perp}>k, & \text { (evanescent modes) }
\end{array}\right.
$$

The propagating modes are important in radiation problems and conventional imaging systems, such as Fourier optics [50]. The evanescent modes are important in the new subject of near-field optics, in which objects can be probed and imaged at nanometer scales improving the resolution of optical microscopy by factors of ten. Some near-field optics references are [177-197].

To prove (D.9), we consider the two-dimensional spatial Fourier transform of $G(\mathbf{r})$ and its inverse. Indicating explicitly the dependence on the coordinates $x, y, z$, we have:

$$
\begin{align*}
g\left(k_{x}, k_{y}, z\right) & =\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} G(x, y, z) e^{j\left(k_{x} x+k_{y} y\right)} d x d y=\frac{e^{-j k_{z}|z|}}{2 j k_{z}}  \tag{D.11}\\
G(x, y, z) & =\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g\left(k_{x}, k_{y}, z\right) e^{-j\left(k_{x} x+k_{y} y\right)} \frac{d k_{x} d k_{y}}{(2 \pi)^{2}}
\end{align*}
$$

Writing $\delta^{(3)}(\mathbf{r})=\delta(x) \delta(y) \delta(z)$ and using the inverse Fourier transform:

$$
\delta(x) \delta(y)=\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} e^{-j\left(k_{x} x+k_{y} y\right)} \frac{d k_{x} d k_{y}}{(2 \pi)^{2}}
$$

we find from Eq. (D.2) that $g\left(k_{x}, k_{y}, z\right)$ must satisfy the one-dimensional Helmholtz Green's function equation (D.3), with $k_{z}^{2}=k^{2}-k_{x}^{2}-k_{y}^{2}=k^{2}-k_{\perp}^{2}$, that is,

$$
\begin{equation*}
\left(\partial_{z}^{2}+k_{z}^{2}\right) g\left(k_{x}, k_{y}, z\right)=-\delta(z) \tag{D.12}
\end{equation*}
$$

whose outgoing/evanescent solution is $g\left(k_{x}, k_{y}, z\right)=e^{-j k_{z}|z|} / 2 j k_{z}$.
A more direct proof of (D.9) is to use cylindrical coordinates, $k_{x}=k_{\perp} \cos \psi, k_{y}=$ $k_{\perp} \sin \psi, x=\rho \cos \phi, y=\rho \sin \phi$, where $k_{\perp}^{2}=k_{x}^{2}+k_{y}^{2}$ and $\rho^{2}=x^{2}+y^{2}$. It follows that
$k_{x} x+k_{y} y=k_{\perp} \rho \cos (\phi-\psi)$. Setting $d x d y=\rho d \rho d \phi=r d r d \phi$, the latter following from $r^{2}=\rho^{2}+z^{2}$, we obtain from Eq. (D.11) after replacing $\rho=\sqrt{r^{2}-z^{2}}$ :

$$
\begin{aligned}
g\left(k_{x}, k_{y}, z\right) & =\iint \frac{e^{-j k r}}{4 \pi r} e^{j\left(k_{x} x+k_{y} y\right)} d x d y=\iint \frac{e^{-j k r}}{4 \pi r} e^{j k_{\perp} \rho \cos (\phi-\psi)} r d r d \phi \\
& =\frac{1}{2} \int_{|z|}^{\infty} d r e^{-j k r} \int_{0}^{2 \pi} \frac{d \phi}{2 \pi} e^{j k_{\perp} \rho \cos (\phi-\psi)}=\frac{1}{2} \int_{|z|}^{\infty} d r e^{-j k r} J_{0}\left(k_{\perp} \sqrt{r^{2}-z^{2}}\right)
\end{aligned}
$$

where we used the integral representation (16.9.2) of the Bessel function $J_{0}(x)$. Looking up the last integral in the table of integrals [104], we find:

$$
\begin{equation*}
g\left(k_{x}, k_{y}, z\right)=\frac{1}{2} \int_{|z|}^{\infty} d r e^{-j k r} J_{0}\left(k_{\perp} \sqrt{r^{2}-z^{2}}\right)=\frac{e^{-j k_{z}|z|}}{2 j k_{z}} \tag{D.13}
\end{equation*}
$$

where $k_{z}$ must be defined exactly as in Eq. (D.10). A direct consequence of Eq. (D.11) and the even-ness of $G(\mathbf{r})$ in $\mathbf{r}$ and of $g\left(k_{x}, k_{y}, z\right)$ in $k_{x}, k_{y}$, is the following result:

$$
\begin{equation*}
\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} e^{-j\left(k_{x} x^{\prime}+k_{y} y^{\prime}\right)} G\left(\mathbf{r}-\mathbf{r}^{\prime}\right) d x^{\prime} d y^{\prime}=e^{-j\left(k_{x} x+k_{y} y\right)} \frac{e^{-j k_{z}\left|z-z^{\prime}\right|}}{2 j k_{z}} \tag{D.14}
\end{equation*}
$$

One can also show the integral:

$$
\int_{0}^{\infty} e^{-j k_{z}^{\prime} z^{\prime}} \frac{e^{-j k_{z}\left|z-z^{\prime}\right|}}{2 j k_{z}} d z^{\prime}= \begin{cases}\frac{e^{-j k_{z}^{\prime} z}}{k_{z}^{\prime 2}-k_{z}^{2}}-\frac{e^{-j k_{z} z}}{2 k_{z}\left(k_{z}^{\prime}-k_{z}\right)}, & \text { for } z \geq 0  \tag{D.15}\\ -\frac{e^{j k_{z} z}}{2 k_{z}\left(k_{z}^{\prime}+k_{z}\right)}, & \text { for } z<0\end{cases}
$$

The proof is obtained by splitting the integral over the sub-intervals [0, $z$ ] and $[z, \infty)$. To handle the limits at infinity, $k_{z}^{\prime}$ must be assumed to be slightly lossy, that is, $k_{z}^{\prime}=\beta_{z}-j \alpha_{z}$, with $\alpha_{z}>0$. Eqs. (D.14) and (D.15) can be combined into:

$$
\int_{V_{+}} e^{-j \boldsymbol{k}^{\prime} \cdot \mathbf{r}^{\prime}} G\left(\mathbf{r}-\mathbf{r}^{\prime}\right) d V^{\prime}= \begin{cases}\frac{e^{-j \boldsymbol{k}^{\prime} \cdot \mathbf{r}}}{k^{\prime 2}-k^{2}}-\frac{e^{-j \boldsymbol{k} \cdot \mathbf{r}}}{2 k_{Z}\left(k_{Z}^{\prime}-k_{Z}\right)}, & \text { for } z \geq 0  \tag{D.16}\\ -\frac{e^{-j \boldsymbol{k}_{-} \cdot \mathbf{r}}}{2 k_{Z}\left(k_{Z}^{\prime}+k_{Z}\right)}, & \text { for } z<0\end{cases}
$$

where $V_{+}$is the half-space $z \geq 0$, and $\boldsymbol{k}, \boldsymbol{k}_{-}, \boldsymbol{k}^{\prime}$ are wave-vectors with the same $k_{x}, k_{y}$ components, but different $k_{z} \mathrm{~s}$ :

$$
\begin{align*}
\boldsymbol{k} & =k_{x} \hat{\mathbf{x}}+k_{y} \hat{\mathbf{y}}+k_{z} \hat{\mathbf{z}} \\
\boldsymbol{k}_{-} & =k_{x} \hat{\mathbf{x}}+k_{y} \hat{\mathbf{y}}-k_{z} \hat{\mathbf{z}}  \tag{D.17}\\
\boldsymbol{k}^{\prime} & =k_{x} \hat{\mathbf{x}}+k_{y} \hat{\mathbf{y}}+k_{z}^{\prime} \hat{\mathbf{z}}
\end{align*}
$$

where we note that $k^{\prime 2}-k^{2}=\left(k_{x}^{2}+k_{y}^{2}+k_{z}^{\prime 2}\right)-\left(k_{x}^{2}+k_{y}^{2}+k_{z}^{2}\right)=k_{z}^{\prime 2}-k_{z}^{2}$.
The Green's function results (D.8)-(D.17) are used in the discussion of the EwaldOseen extinction theorem in Sec. 13.6.

## E. Coordinate Systems

The definitions of cylindrical and spherical coordinates were given in Sec. 13.8. The expressions of the gradient, divergence, curl, Laplacian operators, and delta functions are given below in cartesian, cylindrical, and spherical coordinates.

## Cartesian Coordinates

$$
\begin{aligned}
& \nabla \psi=\hat{\mathbf{x}} \frac{\partial \psi}{\partial x}+\hat{\mathbf{y}} \frac{\partial \psi}{\partial y}+\hat{\mathbf{z}} \frac{\partial \psi}{\partial z} \\
& \nabla^{2} \psi=\frac{\partial^{2} \psi}{\partial x^{2}}+\frac{\partial^{2} \psi}{\partial y^{2}}+\frac{\partial^{2} \psi}{\partial z^{2}} \\
& \nabla \cdot \boldsymbol{A}=\frac{\partial A_{x}}{\partial x}+\frac{\partial A_{y}}{\partial y}+\frac{\partial A_{z}}{\partial z} \\
& \nabla \times \boldsymbol{A}=\hat{\mathbf{x}}\left(\frac{\partial A_{z}}{\partial y}-\frac{\partial A_{y}}{\partial z}\right)+\hat{\mathbf{y}}\left(\frac{\partial A_{x}}{\partial z}-\frac{\partial A_{z}}{\partial x}\right)+\hat{\mathbf{z}}\left(\frac{\partial A_{y}}{\partial x}-\frac{\partial A_{x}}{\partial y}\right) \\
&=\left|\begin{array}{ccc}
\hat{\mathbf{x}} & \hat{\mathbf{y}} & \hat{\mathbf{z}} \\
\frac{\partial}{\partial_{x}} & \frac{\partial}{\partial y} & \frac{\partial}{\partial_{z}} \\
A_{x} & A_{y} & A_{z}
\end{array}\right| \\
& \delta^{(3)}\left(\mathbf{r}-\mathbf{r}^{\prime}\right)=\delta\left(x-x^{\prime}\right) \delta\left(y-y^{\prime}\right) \delta\left(z-z^{\prime}\right)
\end{aligned}
$$

## Cylindrical Coordinates

$$
\begin{align*}
& \nabla \psi=\hat{\boldsymbol{\rho}} \frac{\partial \psi}{\partial \rho}+\hat{\boldsymbol{\phi}} \frac{1}{\rho} \frac{\partial \psi}{\partial \phi}+\hat{\mathbf{z}} \frac{\partial \psi}{\partial z}  \tag{E.2a}\\
& \nabla^{2} \psi=\frac{1}{\rho} \frac{\partial}{\partial \rho}\left(\rho \frac{\partial \psi}{\partial \rho}\right)+\frac{1}{\rho^{2}} \frac{\partial^{2} \psi}{\partial \phi^{2}}+\frac{\partial^{2} \psi}{\partial z^{2}} \\
& \nabla \cdot \boldsymbol{A}=\frac{1}{\rho} \frac{\partial\left(\rho A_{\rho}\right)}{\partial \rho}+\frac{1}{\rho} \frac{\partial A_{\phi}}{\partial \phi}+\frac{\partial A_{z}}{\partial z}  \tag{E.2c}\\
& \nabla \times \boldsymbol{A}=\hat{\boldsymbol{\rho}}\left(\frac{1}{\rho} \frac{\partial A_{z}}{\partial \phi}-\frac{\partial A_{\phi}}{\partial z}\right)+\hat{\boldsymbol{\phi}}\left(\frac{\partial A_{\rho}}{\partial z}-\frac{\partial A_{z}}{\partial \rho}\right)+\hat{\mathbf{z}} \frac{1}{\rho}\left(\frac{\partial\left(\rho A_{\phi}\right)}{\partial \rho}-\frac{\partial A_{\rho}}{\partial \phi}\right) \\
& \delta^{(3)}\left(\mathbf{r}-\mathbf{r}^{\prime}\right)=\frac{1}{\rho} \delta\left(\rho-\rho^{\prime}\right) \delta\left(\phi-\phi^{\prime}\right) \delta\left(z-z^{\prime}\right)
\end{align*}
$$

## Spherical Coordinates

$$
\begin{aligned}
\nabla \psi & =\hat{\mathbf{r}} \frac{\partial \psi}{\partial r}+\hat{\boldsymbol{\theta}} \frac{1}{r} \frac{\partial \psi}{\partial \theta}+\hat{\boldsymbol{\phi}} \frac{1}{r \sin \theta} \frac{\partial \psi}{\partial \phi} \\
\nabla^{2} \psi & =\frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} \frac{\partial \psi}{\partial r}\right)+\frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial \psi}{\partial \theta}\right)+\frac{1}{r^{2} \sin ^{2} \theta} \frac{\partial^{2} \psi}{\partial \phi^{2}}
\end{aligned}
$$

$$
\begin{align*}
& \boldsymbol{\nabla} \cdot \boldsymbol{A}=\frac{1}{r^{2}} \frac{\partial\left(r^{2} A_{r}\right)}{\partial r}+\frac{1}{r \sin \theta} \frac{\partial\left(\sin \theta A_{\theta}\right)}{\partial \theta}+\frac{1}{r \sin \theta} \frac{\partial A_{\phi}}{\partial \phi}  \tag{E.3c}\\
& \nabla \times \boldsymbol{A}=\hat{\mathbf{r}} \frac{1}{r \sin \theta}\left(\frac{\partial\left(\sin \theta A_{\phi}\right)}{\partial \theta}-\frac{\partial A_{\theta}}{\partial \phi}\right)+\hat{\boldsymbol{\theta}} \frac{1}{r}\left(\frac{1}{\sin \theta} \frac{\partial A_{r}}{\partial \phi}-\frac{\partial\left(r A_{\phi}\right)}{\partial r}\right)  \tag{E.3d}\\
& +\hat{\boldsymbol{\phi}} \frac{1}{r}\left(\frac{\partial\left(r A_{\theta}\right)}{\partial r}-\frac{\partial A_{r}}{\partial \theta}\right) \\
& \delta^{(3)}\left(\mathbf{r}-\mathbf{r}^{\prime}\right)=\frac{1}{r^{2} \sin \theta} \delta\left(r-r^{\prime}\right) \delta\left(\theta-\theta^{\prime}\right) \delta\left(\phi-\phi^{\prime}\right) \tag{E.3e}
\end{align*}
$$

## Transformations Between Coordinate Systems

A vector $\boldsymbol{A}$ can be expressed component-wise in the three coordinate systems as:

$$
\begin{align*}
\boldsymbol{A} & =\hat{\mathbf{x}} A_{x}+\hat{\mathbf{y}} A_{y}+\hat{\mathbf{z}} A_{z} \\
& =\hat{\boldsymbol{\rho}} A_{\rho}+\hat{\boldsymbol{\phi}} A_{\phi}+\hat{\mathbf{z}} A_{z}  \tag{E.4}\\
& =\hat{\mathbf{r}} A_{r}+\hat{\boldsymbol{\theta}} A_{\theta}+\hat{\boldsymbol{\phi}} A_{\phi}
\end{align*}
$$

The components in one coordinate system can be expressed in terms of the components of another by using the following relationships between the unit vectors, which were also given in Eqs. (13.8.1)-(13.8.3):

$$
\begin{array}{clr}
x=\rho \cos \phi & \hat{\boldsymbol{\rho}}=\hat{\mathbf{x}} \cos \phi+\hat{\mathbf{y}} \sin \phi & \hat{\mathbf{x}}=\hat{\boldsymbol{\rho}} \cos \phi-\hat{\boldsymbol{\phi}} \sin \phi \\
y=\rho \sin \phi & \hat{\boldsymbol{\phi}}=-\hat{\mathbf{x}} \sin \phi+\hat{\mathbf{y}} \cos \phi & \hat{\mathbf{y}}=\hat{\boldsymbol{\rho}} \sin \phi+\hat{\boldsymbol{\phi}} \cos \phi \\
\rho=r \sin \theta & \hat{\mathbf{r}}=\hat{\mathbf{z}} \cos \theta+\hat{\boldsymbol{\rho}} \sin \theta & \hat{\mathbf{z}}=\hat{\mathbf{r}} \cos \theta-\hat{\boldsymbol{\theta}} \sin \theta \\
z=r \cos \theta & \hat{\boldsymbol{\theta}}=-\hat{\mathbf{z}} \sin \theta+\hat{\boldsymbol{\rho}} \cos \theta & \hat{\boldsymbol{\rho}}=\hat{\mathbf{r}} \sin \theta+\hat{\boldsymbol{\theta}} \cos \theta \\
& \\
& x=r \sin \theta \cos \phi & \hat{\mathbf{r}}=\hat{\mathbf{x}} \cos \phi \sin \theta+\hat{\mathbf{y}} \sin \phi \sin \theta+\hat{\mathbf{z}} \cos \theta  \tag{E.7}\\
y=r \sin \theta \sin \phi & \hat{\boldsymbol{\theta}}=\hat{\mathbf{x}} \cos \phi \cos \theta+\hat{\mathbf{y}} \sin \phi \cos \theta-\hat{\mathbf{z}} \sin \theta \\
z=r \cos \theta & \hat{\boldsymbol{\phi}}=-\hat{\mathbf{x}} \sin \phi+\hat{\mathbf{y}} \cos \phi
\end{array}
$$

$$
\begin{align*}
& \hat{\mathbf{x}}=\hat{\mathbf{r}} \sin \theta \cos \phi+\hat{\boldsymbol{\theta}} \cos \theta \cos \phi-\hat{\boldsymbol{\phi}} \sin \phi \\
& \hat{\mathbf{y}}=\hat{\mathbf{r}} \sin \theta \sin \phi+\hat{\boldsymbol{\theta}} \cos \theta \sin \phi+\hat{\boldsymbol{\phi}} \cos \phi  \tag{E.8}\\
& \hat{\mathbf{z}}=\hat{\mathbf{r}} \cos \theta-\hat{\boldsymbol{\theta}} \sin \theta
\end{align*}
$$

For example, to express the spherical components $A_{\theta}, A_{\phi}$ in terms of the cartesian components, we proceed as follows:

$$
\begin{aligned}
& A_{\theta}=\hat{\boldsymbol{\theta}} \cdot \boldsymbol{A}=\hat{\boldsymbol{\theta}} \cdot\left(\hat{\mathbf{x}} A_{x}+\hat{\mathbf{y}} A_{y}+\hat{\mathbf{z}} A_{z}\right)=(\hat{\boldsymbol{\theta}} \cdot \hat{\mathbf{x}}) A_{x}+(\hat{\boldsymbol{\theta}} \cdot \hat{\mathbf{y}}) A_{y}+(\hat{\boldsymbol{\theta}} \cdot \hat{\mathbf{z}}) A_{z} \\
& A_{\phi}=\hat{\boldsymbol{\phi}} \cdot \boldsymbol{A}=\hat{\boldsymbol{\phi}} \cdot\left(\hat{\mathbf{x}} A_{x}+\hat{\mathbf{y}} A_{y}+\hat{\mathbf{z}} A_{z}\right)=(\hat{\boldsymbol{\phi}} \cdot \hat{\mathbf{x}}) A_{x}+(\hat{\boldsymbol{\phi}} \cdot \hat{\mathbf{y}}) A_{y}+(\hat{\boldsymbol{\phi}} \cdot \hat{\mathbf{z}}) A_{z}
\end{aligned}
$$

The dot products can be read off Eq. (E.7), resulting in:

$$
\begin{align*}
& A_{\theta}=\cos \phi \cos \theta A_{x}+\sin \phi \cos \theta A_{y}-\sin \theta A_{z}  \tag{E.9}\\
& A_{\phi}=-\sin \phi A_{x}+\cos \phi A_{y}
\end{align*}
$$

Similarly, using Eq. (E.6) the cylindrical components $A_{\rho}, A_{z}$ can be expressed in terms of spherical components as:

$$
\begin{align*}
& A_{\rho}=\hat{\boldsymbol{\rho}} \cdot \boldsymbol{A}=\hat{\boldsymbol{\rho}} \cdot\left(\hat{\mathbf{r}} A_{r}+\hat{\boldsymbol{\theta}} A_{\theta}+\hat{\boldsymbol{\phi}} A_{\phi}\right)=\sin \theta A_{r}+\cos \theta A_{\theta} \\
& A_{z}=\hat{\mathbf{z}} \cdot \boldsymbol{A}=\hat{\mathbf{z}} \cdot\left(\hat{\mathbf{r}} A_{r}+\hat{\boldsymbol{\theta}} A_{\theta}+\hat{\boldsymbol{\phi}} A_{\phi}\right)=\cos \theta A_{r}-\cos \theta A_{\theta} \tag{E.10}
\end{align*}
$$

## F. Fresnel Integrals

The Fresnel functions $C(x)$ and $S(x)$ are defined by [103]:

$$
\begin{equation*}
C(x)=\int_{0}^{x} \cos \left(\frac{\pi}{2} t^{2}\right) d t, \quad S(x)=\int_{0}^{x} \sin \left(\frac{\pi}{2} t^{2}\right) d t \tag{F.1}
\end{equation*}
$$

They may be combined into the complex function:

$$
\begin{equation*}
F(x)=C(x)-j S(x)=\int_{0}^{x} e^{-j(\pi / 2) t^{2}} d t \tag{F.2}
\end{equation*}
$$

$C(x), S(x)$, and $F(x)$ are odd functions of $x$ and have the asymptotic values:

$$
\begin{equation*}
C(\infty)=S(\infty)=\frac{1}{2}, \quad F(\infty)=\frac{1-j}{2} \tag{F.3}
\end{equation*}
$$

At $x=0$, we have $F(0)=0$ and $F^{\prime}(0)=1$, so that the Taylor series approximation is $F(x) \simeq x$, for small $x$. The asymptotic expansions of $C(x), S(x)$, and $F(x)$ are for large positive $x$ :

$$
\begin{align*}
& F(x)=\frac{1-j}{2}+\frac{j}{\pi x} e^{-j \pi x^{2} / 2} \\
& C(x)=\frac{1}{2}+\frac{1}{\pi x} \sin \left(\frac{\pi}{2} x^{2}\right)  \tag{F.4}\\
& S(x)=\frac{1}{2}-\frac{1}{\pi x} \cos \left(\frac{\pi}{2} x^{2}\right)
\end{align*}
$$

Associated with $C(x)$ and $S(x)$ are the type-2 Fresnel integrals:

$$
\begin{equation*}
C_{2}(x)=\int_{0}^{x} \frac{\cos t}{\sqrt{2 \pi t}} d t, \quad S_{2}(x)=\int_{0}^{x} \frac{\sin t}{\sqrt{2 \pi t}} d t \tag{F.5}
\end{equation*}
$$

They are combined into the complex function:

$$
\begin{equation*}
F_{2}(x)=C_{2}(x)-j S_{2}(x)=\int_{0}^{x} \frac{e^{-j t}}{\sqrt{2 \pi t}} d t \tag{F.6}
\end{equation*}
$$

The two types are related by, if $x \geq 0$ :

$$
\begin{equation*}
C(x)=C_{2}\left(\frac{\pi}{2} x^{2}\right), \quad S(x)=S_{2}\left(\frac{\pi}{2} x^{2}\right), \quad F(x)=F_{2}\left(\frac{\pi}{2} x^{2}\right) \tag{F.7}
\end{equation*}
$$

and if $x<0$, we set $F(x)=-F(-x)=-F_{2}\left(\pi x^{2} / 2\right)$.

The Fresnel function $F_{2}(x)$ can be evaluated numerically using Boersma's approximation [729], which achieves a maximum error of $10^{-9}$ over all $x$. The algorithm approximates the function $F_{2}(x)$ as follows:

$$
F_{2}(x)= \begin{cases}e^{-j x} \sqrt{\frac{x}{4}} \sum_{n=0}^{11}\left(a_{n}+j b_{n}\right)\left(\frac{x}{4}\right)^{n}, & \text { if } \quad 0 \leq x \leq 4  \tag{F.8}\\ \frac{1-j}{2}+e^{-j x} \sqrt{\frac{4}{x}} \sum_{n=0}^{11}\left(c_{n}+j d_{n}\right)\left(\frac{4}{x}\right)^{n}, & \text { if } \quad x>4\end{cases}
$$

where the coefficients $a_{n}, b_{n}, c_{n}, d_{n}$ are given in [729]. Consistency with the small- and large- $x$ expansions of $F(x)$ requires that $a_{0}+j b_{0}=\sqrt{8 / \pi}$ and $c_{0}+j d_{0}=j / \sqrt{8 \pi}$. We have implemented Eq. (F.8) with the MATLAB function fcs2:
$\mathrm{F} 2=\mathrm{fcs} 2(\mathrm{x}) ; \quad \%$ Fresnel integrals $F \mathrm{~b} 2(\mathrm{x})=C \mathrm{~b} 2(x)-j \mathrm{Sb} 2(x)$
The ordinary Fresnel integral $F(x)$ can be computed with the help of Eq. (F.7). The MATLAB function fcs calculates $F(x)$ for any vector of values $x$ by calling fcs2:

$$
\mathrm{F}=\mathrm{fcs}(\mathrm{x}) ; \quad \% \text { Fresnel integrals } F(x)=C(x)-j S(x)
$$

In calculating the radiation patterns of pyramidal horns, it is desired to calculate a Fresnel diffraction integral of the type:

$$
\begin{equation*}
F_{0}(\nu, \sigma)=\int_{-1}^{1} e^{j \pi \nu \xi} e^{-j(\pi / 2) \sigma^{2} \xi^{2}} d \xi \tag{F.9}
\end{equation*}
$$

Making the variable change $t=\sigma \xi-v / \sigma$, this integral can be computed in terms of the Fresnel function $F(x)=C(x)-j S(x)$ as follows:

$$
\begin{equation*}
F_{0}(\nu, \sigma)=\frac{1}{\sigma} e^{j(\pi / 2)\left(\nu^{2} / \sigma^{2}\right)}\left[F\left(\frac{v}{\sigma}+\sigma\right)-F\left(\frac{v}{\sigma}-\sigma\right)\right] \tag{F.10}
\end{equation*}
$$

where we also used the oddness of $F(x)$. The value of Eq. (F.9) at $v=0$ is:

$$
\begin{equation*}
F_{0}(0, \sigma)=\frac{1}{\sigma}[F(\sigma)-F(-\sigma)]=2 \frac{F(\sigma)}{\sigma} \tag{F.11}
\end{equation*}
$$

Eq. (F.10) assumes that $\sigma \neq 0$. If $\sigma=0$, the integral (F.9) reduces to the sinc function:

$$
\begin{equation*}
F_{0}(v, 0)=2 \frac{\sin (\pi v)}{\pi v} \tag{F.12}
\end{equation*}
$$

From either (F.11) or (F.12), we find $F_{0}(0,0)=2$. A related integral that is also required in the theory of horns is the following:

$$
\begin{equation*}
F_{1}(v, \sigma)=\int_{-1}^{1} \cos \left(\frac{\pi \xi}{2}\right) e^{j \pi v \xi} e^{-j(\pi / 2) \sigma^{2} \xi^{2}} d \xi \tag{F.13}
\end{equation*}
$$

Writing $\cos (\pi \xi / 2)=\left(e^{j \pi \xi / 2}+e^{-j \pi \xi / 2}\right) / 2$, the integral $F_{1}(\nu, s)$ can be expressed in terms of $F_{0}(\nu, \sigma)$ as follows:

$$
\begin{equation*}
F_{1}(\nu, \sigma)=\frac{1}{2}\left[F_{0}(v+0.5, \sigma)+F_{0}(v-0.5, \sigma)\right] \tag{F.14}
\end{equation*}
$$

It can be verified easily that $F_{0}(0.5, \sigma)=F_{0}(-0.5, \sigma)$, therefore, the value of $F_{1}(\nu, \sigma)$ at $v=0$ will be given by:

$$
\begin{equation*}
F_{1}(0, \sigma)=F_{0}(0.5, \sigma)=\frac{1}{\sigma} e^{j \pi /\left(8 \sigma^{2}\right)}\left[F\left(\frac{1}{2 \sigma}+\sigma\right)-F\left(\frac{1}{2 \sigma}-\sigma\right)\right] \tag{F.15}
\end{equation*}
$$

Using the asymptotic expansion (F.4), we find the expansion valid for small $\sigma$ :

$$
\begin{equation*}
F\left(\frac{1}{2 \sigma} \pm \sigma\right)=\frac{1-j}{2} \mp \frac{2 \sigma}{\pi} e^{-j \pi /\left(8 \sigma^{2}\right)}, \quad \text { for small } \sigma \tag{F.16}
\end{equation*}
$$

For $\sigma=0$, the integral $F_{1}(\nu, \sigma)$ reduces to the double-sinc function:

$$
\begin{array}{r}
F_{1}(v, 0)=\int_{-1}^{1} \cos \left(\frac{\pi \xi}{2}\right) e^{j \pi v \xi} d \xi=\frac{1}{2}\left[F_{0}(v+0.5,0)+F_{0}(v-0.5,0)\right]  \tag{F.17}\\
=\frac{\sin (\pi(v+0.5))}{\pi(v+0.5)}+\frac{\sin (\pi(v-0.5))}{\pi(v-0.5)}=\frac{4}{\pi} \frac{\cos (\pi v)}{1-4 v^{2}}
\end{array}
$$

From either Eq. (F.16) or (F.17), we find $F_{1}(0,0)=4 / \pi$.
The MATLAB function diffint can be used to evaluate both Eq. (F.9) and (F.13) for any vector of values $v$ and any vector of positive numbers $\sigma$, including $\sigma=0$. It calls fcs to evaluate the diffraction integral (F.9) according to Eq. (F.10). Its usage is:

$$
\begin{array}{ll}
\mathrm{F} 0=\text { diffint }(\mathrm{v}, \text { sigma }, 0) ; & \text { \% diffraction integral } F \mathrm{~b} 0(v, \sigma) \text {, Eq. (F.9) } \\
\text { F1 }=\text { diffint }(v, \text { sigma }, 1) ; & \text { \% diffraction integral } F \mathrm{~b} 1(v, \sigma) \text {, Eq. (F.13) }
\end{array}
$$

The vectors $v$, sigma can be entered either as rows or columns, but the result will be a matrix of size length(v) x length(sigma). The integral $F_{0}(v, \sigma)$ can also be calculated by the simplified call:

$$
\text { F0 }=\text { diffint }(v, \text { sigma })
$$

\% diffraction integral $F \mathrm{~b} 0(v, \sigma)$, Eq. (F.9)
Actually, the most general syntax of diffint is as follows:
$\mathrm{F}=\operatorname{diffint}(\mathrm{v}$, sigma, $\mathrm{a}, \mathrm{c} 1, \mathrm{c} 2) ; \quad$ \% diffraction integral $F(v, \sigma, a)$, Eq. (F.18)
It evaluates the more general integral:

$$
\begin{equation*}
F(v, \sigma, a)=\int_{c_{1}}^{c_{2}} \cos \left(\frac{\pi \xi a}{2}\right) e^{j \pi v \xi} e^{-j(\pi / 2) \sigma^{2} \xi^{2}} d \xi \tag{F.18}
\end{equation*}
$$

For $a=0$, we have:

$$
\begin{equation*}
F(v, \sigma, 0)=\frac{1}{\sigma} e^{j(\pi / 2)\left(v^{2} / \sigma^{2}\right)}\left[F\left(\frac{v}{\sigma}-\sigma c_{1}\right)-F\left(\frac{v}{\sigma}-\sigma c_{2}\right)\right] \tag{F.19}
\end{equation*}
$$

For $a \neq 0$, we can express $F(\nu, \sigma, a)$ in terms of $F(\nu, \sigma, 0)$ :

$$
F(v, \sigma, a)=\frac{1}{2}[F(v+0.5 a, \sigma, 0)+F(v-0.5 a, \sigma, 0)]
$$

For $a=0$ and $\sigma=0, F(v, \sigma, a)$ reduces to the complex sinc function:

$$
F(v, 0,0)=\frac{e^{j \pi v c_{2}}-e^{j \pi v c_{1}}}{j \pi v}=\left(c_{2}-c_{1}\right) \frac{\sin \left(\pi\left(c_{2}-c_{1}\right) v / 2\right)}{\pi\left(c_{2}-c_{1}\right) v / 2} e^{j \pi\left(c_{2}+c_{1}\right) v / 2}
$$

## G. Lorentz Transformations

According to Einstein's special theory of relativity [123], Lorentz transformations describe the transformation between the space-time coordinates of two coordinate systems moving relative to each other at constant velocity. Maxwell's equations remain invariant under Lorentz transformations. This is demonstrated below.

Let the two coordinate frames be $S$ and $S^{\prime}$. By convention, we may think of $S$ as the "fixed" laboratory frame with respect to which the frame $S^{\prime}$ is moving at a constant velocity $\boldsymbol{v}$. For example, if $\boldsymbol{v}$ is in the $z$-direction, the space-time coordinates $\{t, x, y, z\}$ of $S$ are related to the coordinates $\left\{t^{\prime}, x^{\prime}, y^{\prime}, z^{\prime}\right\}$ of $S^{\prime}$ by the Lorentz transformation:

| $t^{\prime}$ | $=\gamma\left(t-\frac{v}{c^{2}} z\right)$ |
| ---: | :--- |
| $z^{\prime}$ | $=\gamma(z-v t)$ |
| $x^{\prime}$ | $=x$ |
| $y^{\prime}$ | $=y$ |

$$
\text { where } \quad \gamma=\frac{1}{\sqrt{1-v^{2} / c^{2}}}
$$


where $c$ is the speed of light in vacuum. Defining the scaled quantities $\tau=c t$ and $\beta=v / c$, the above transformation and its inverse, obtained by replacing $\beta$ by $-\beta$, may be written as follows:

$$
\begin{align*}
& \tau^{\prime}=\gamma(\tau-\beta z) \\
& z^{\prime}=\gamma(z-\beta \tau) \\
& x^{\prime}=x \\
& y^{\prime}=y
\end{align*} \Leftrightarrow\left[\begin{array}{l}
\tau=\gamma\left(\tau^{\prime}+\beta z^{\prime}\right) \\
z=\gamma\left(z^{\prime}+\beta \tau^{\prime}\right)  \tag{G.1}\\
x=x^{\prime} \\
y=y^{\prime}
\end{array}\right.
$$

These transformations are also referred to as Lorentz boosts to indicate the fact that one frame is boosted to move relative to the other. Interchanging the roles of $z$ and $x$, or $z$ and $y$, one obtains the Lorentz transformations for motion along the $x$ or $y$ directions, respectively. Eqs. (G.1) may be expressed more compactly in matrix form:

$$
\mathrm{x}^{\prime}=L \mathrm{x} \text {, where } \mathrm{x}=\left[\begin{array}{c}
\tau  \tag{G.2}\\
x \\
y \\
z
\end{array}\right], \quad \mathrm{x}^{\prime}=\left[\begin{array}{c}
\tau^{\prime} \\
x^{\prime} \\
y^{\prime} \\
z^{\prime}
\end{array}\right], \quad L=\left[\begin{array}{cccc}
\gamma & 0 & 0 & -\gamma \beta \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
-\gamma \beta & 0 & 0 & \gamma
\end{array}\right]
$$

Such transformations leave the quadratic form $\left(c^{2} t^{2}-x^{2}-y^{2}-z^{2}\right)$ invariant, that is,

$$
\begin{equation*}
c^{2} t^{\prime 2}-x^{\prime 2}-y^{\prime 2}-z^{\prime 2}=c^{2} t^{2}-x^{2}-y^{2}-z^{2} \tag{G.3}
\end{equation*}
$$

Introducing the diagonal metric matrix $G=\operatorname{diag}(1,-1,-1,-1)$, we may write the quadratic form as follows, where $\mathrm{x}^{T}$ denotes the transposed vector, that is, the row vector $\mathrm{x}^{T}=[\tau, x, y, z]$ :

$$
\begin{equation*}
\mathrm{x}^{T} G \mathrm{x}=\tau^{2}-x^{2}-y^{2}-z^{2}=c^{2} t^{2}-x^{2}-y^{2}-z^{2} \tag{G.4}
\end{equation*}
$$

More generally, a Lorentz transformation is defined as any linear transformation $\mathrm{x}^{\prime}=$ $L x$ that leaves the quadratic form $x^{T} G x$ invariant. The invariance condition requires that: $\mathrm{x}^{\prime T} G \mathrm{x}^{\prime}=\mathrm{x}^{T} L^{T} G L \mathrm{x}=\mathrm{x}^{T} G \mathrm{x}$, or

$$
\begin{equation*}
L^{T} G L=G \tag{G.5}
\end{equation*}
$$

In addition to the Lorentz boosts of Eq. (G.1), the more general transformations satisfying (G.5) include rotations of the three spatial coordinates, as well as time or space reflections. For example, a rotation has the form:

$$
L=\left[\begin{array}{l|lll}
1 & 0 & 0 & 0 \\
\hline 0 & & & \\
0 & & R & \\
0 & & &
\end{array}\right]
$$

where $R$ is a $3 \times 3$ orthogonal rotation matrix, that is, $R^{T} R=I$, where $I$ is the $3 \times 3$ identity matrix. The most general Lorentz boost corresponding to arbitrary velocity $\boldsymbol{v}=\left[v_{x}, v_{y}, v_{z}\right]^{T}$ is given by:

$$
L=\left[\begin{array}{c|c}
\gamma & -\gamma \boldsymbol{\beta}^{T}  \tag{G.6}\\
\hline-\gamma \boldsymbol{\beta} & I+\frac{\gamma^{2}}{\gamma+1} \boldsymbol{\beta} \boldsymbol{\beta}^{T}
\end{array}\right], \quad \text { where } \boldsymbol{\beta}=\frac{\boldsymbol{v}}{c}, \quad \gamma=\frac{1}{\sqrt{1-\boldsymbol{\beta}^{T} \boldsymbol{\beta}}}
$$

When $\boldsymbol{v}=[0,0, v]^{T}$, or $\boldsymbol{\beta}=[0,0, \beta]^{T}$, Eq. (G.6) reduces to (G.1). Defining $\beta=|\boldsymbol{\beta}|=$ $\sqrt{\boldsymbol{\beta}^{T} \boldsymbol{\beta}}$ and the unit vector $\hat{\boldsymbol{\beta}}=\boldsymbol{\beta} / \beta$, and using the relationship $\gamma^{2} \beta^{2}=\gamma^{2}-1$, it can be verified that the spatial part of the matrix $L$ can be written in the form:

$$
\begin{equation*}
I+\frac{\gamma^{2}}{\gamma+1} \boldsymbol{\beta} \boldsymbol{\beta}^{T}=I+(\gamma-1) \hat{\boldsymbol{\beta}} \hat{\boldsymbol{\beta}}^{T} \tag{G.7}
\end{equation*}
$$

The set of matrices $L$ satisfying Eq. (G.5) forms a group called the Lorentz group. In particular, the $z$-directed boosts of Eq. (G.2) form a commutative subgroup. Denoting these boosts by $L(\beta)$, the application of two successive boosts by velocity factors $\beta_{1}=$ $\nu_{1} / \mathcal{C}$ and $\beta_{2}=\nu_{2} / c$ leads to the combined boost $L(\beta)=L\left(\beta_{1}\right) L\left(\beta_{2}\right)$, where:

$$
\begin{equation*}
\beta=\frac{\beta_{1}+\beta_{2}}{1+\beta_{1} \beta_{2}} \quad \Leftrightarrow \quad v=\frac{v_{1}+v_{2}}{1+v_{1} v_{2} / c^{2}} \tag{G.8}
\end{equation*}
$$

with $\beta=v / c$. Eq. (G.8) is Einstein's relativistic velocity addition theorem. The same group property implies also that $L^{-1}(\beta)=L(-\beta)$. The proof of Eq. (G.8) follows from the following condition, where $\gamma_{1}=1 / \sqrt{1-\beta_{1}^{2}}$ and $\gamma_{2}=1 / \sqrt{1-\beta_{2}^{2}}$ :
$\left[\begin{array}{cccc}\gamma & 0 & 0 & -\gamma \beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma \beta & 0 & 0 & \gamma\end{array}\right]=\left[\begin{array}{cccc}\gamma_{1} & 0 & 0 & -\gamma_{1} \beta_{1} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma_{1} \beta_{1} & 0 & 0 & \gamma_{1}\end{array}\right]\left[\begin{array}{cccc}\gamma_{2} & 0 & 0 & -\gamma_{2} \beta_{2} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma_{2} \beta_{2} & 0 & 0 & \gamma_{2}\end{array}\right]$

A four-vector is a four-dimensional vector that transforms like the vector x under Lorentz transformations, that is, its components with respect to the two moving frames $S$ and $S^{\prime}$ are related by:

$$
a^{\prime}=L a, \quad \text { where } \quad a=\left[\begin{array}{l}
a_{0}  \tag{G.9}\\
a_{x} \\
a_{y} \\
a_{z}
\end{array}\right], \quad a^{\prime}=\left[\begin{array}{l}
a_{0}^{\prime} \\
a_{x}^{\prime} \\
a_{y}^{\prime} \\
a_{z}^{\prime}
\end{array}\right]
$$

For example, under the $z$-directed boost of Eq. (G.1), the four-vector $a$ will transform as:

$$
\begin{align*}
& a_{0}^{\prime}=\gamma\left(a_{0}-\beta a_{z}\right)  \tag{G.10}\\
& a_{z}^{\prime}=\gamma\left(a_{z}-\beta a_{0}\right) \\
& a_{x}^{\prime}=a_{x} \\
& a_{y}^{\prime}=a_{y}
\end{aligned} \quad \Leftrightarrow \quad \begin{aligned}
& a_{0}=\gamma\left(a_{0}^{\prime}+\beta a_{z}^{\prime}\right) \\
& a_{z}=\gamma\left(a_{z}^{\prime}+\beta a_{0}^{\prime}\right) \\
& a_{x}=a_{x}^{\prime} \\
& a_{y}^{\prime}=a_{y}^{\prime}
\end{align*}
$$

Four-vectors transforming according to Eq. (G.9) are referred to as contravariant. Under the general Lorentz boost of Eq. (G.6), the spatial components of $a$ that are transverse to the direction of the velocity vector $v$ remain unchanged, whereas the parallel component transforms as in Eq. (G.10), that is, the most general Lorentz boost transformation for a four-vector takes the form:

$$
\begin{align*}
& a_{0}^{\prime}=\gamma\left(a_{0}-\beta a_{\|}\right)  \tag{G.11}\\
& a_{\|}^{\prime}=\gamma\left(a_{\|}-\beta a_{0}\right) \\
& \boldsymbol{a}_{\perp}^{\prime}=\boldsymbol{a}_{\perp}
\end{align*} \quad \gamma=\frac{1}{\sqrt{1-\beta^{2}}}, \quad \beta=|\boldsymbol{\beta}|, \quad \boldsymbol{\beta}=\frac{\boldsymbol{v}}{\boldsymbol{c}}
$$

where $a_{\|}=\hat{\boldsymbol{\beta}}^{T} \boldsymbol{a}$ and $\boldsymbol{a}=\left[a_{x}, a_{y}, a_{z}\right]^{T}$ is the spatial part of $a$. Then,

$$
\boldsymbol{a}_{\|}=\hat{\boldsymbol{\beta}} a_{\|}=\hat{\boldsymbol{\beta}}\left(\hat{\boldsymbol{\beta}}^{T} \boldsymbol{a}\right) \quad \text { and } \quad \boldsymbol{a}_{\perp}=\boldsymbol{a}-\boldsymbol{a}_{\|}=\boldsymbol{a}-\hat{\boldsymbol{\beta}} a_{\|}
$$

Setting $\boldsymbol{\beta}=\beta \hat{\boldsymbol{\beta}}$ and using Eq. (G.7), the Lorentz transformation (G.6) gives:

$$
\left[\begin{array}{c}
a_{0}^{\prime} \\
\boldsymbol{a}^{\prime}
\end{array}\right]=\left[\begin{array}{cc}
\gamma & -\gamma \beta \hat{\boldsymbol{\beta}}^{T} \\
-\gamma \beta \hat{\boldsymbol{\beta}} & I+(\gamma-1) \hat{\boldsymbol{\beta}} \hat{\boldsymbol{\beta}}^{T}
\end{array}\right]\left[\begin{array}{c}
a_{0} \\
\boldsymbol{a}
\end{array}\right]=\left[\begin{array}{c}
\gamma\left(a_{0}-\beta a_{\|}\right) \\
\boldsymbol{a}-\hat{\boldsymbol{\beta}} a_{\|}+\hat{\boldsymbol{\beta}} \gamma\left(a_{\|}-\beta a_{0}\right)
\end{array}\right]
$$

from which Eq. (G.11) follows.
For any two four-vectors $a, b$, the quadratic form $a^{T} G b$ remains invariant under Lorentz transformations, that is, $a^{\prime T} G b^{\prime}=a^{T} G b$, or,

$$
a_{0}^{\prime} b_{0}^{\prime}-\boldsymbol{a}^{\prime} \cdot \boldsymbol{b}^{\prime}=a_{0} b_{0}-\boldsymbol{a} \cdot \boldsymbol{b}, \quad \text { where } \quad a=\left[\begin{array}{c}
a_{0}  \tag{G.12}\\
\boldsymbol{a}
\end{array}\right], \quad b=\left[\begin{array}{c}
b_{0} \\
\boldsymbol{b}
\end{array}\right]
$$

Some examples of four-vectors are given in the following table:

| four-vector | $a_{0}$ | $a_{x}$ | $a_{y}$ | $a_{z}$ |
| :--- | :---: | :---: | :---: | :---: |
| time and space | $c t$ | $x$ | $y$ | $z$ |
| frequency and wavenumber | $\omega / c$ | $k_{x}$ | $k_{y}$ | $k_{z}$ |
| energy and momentum | $E / c$ | $p_{x}$ | $p_{y}$ | $p_{z}$ |
| charge and current densities | $c \rho$ | $J_{x}$ | $J_{y}$ | $J_{z}$ |
| scalar and vector potentials | $\varphi$ | $c A_{x}$ | $c A_{y}$ | $c A_{z}$ |

For example, under the $z$-directed boost of Eq. (G.1), the frequency-wavenumber transformation will be as follows:

$$
\left.\begin{align*}
& \omega^{\prime}=\gamma\left(\omega-\beta c k_{z}\right) \\
& k_{z}^{\prime}=\gamma\left(k_{z}-\frac{\beta}{c} \omega\right) \\
& k_{x}^{\prime}=k_{x} \\
& k_{y}^{\prime}=k_{y}
\end{aligned} \Leftrightarrow \Leftrightarrow \begin{aligned}
& \omega=\gamma\left(\omega^{\prime}+\beta c k_{z}^{\prime}\right) \\
& k_{z}=\gamma\left(k_{z}^{\prime}+\frac{\beta}{c} \omega^{\prime}\right)  \tag{G.14}\\
& k_{x}=k_{x}^{\prime} \\
& k_{y}=k_{y}^{\prime}
\end{align*} \right\rvert\,, \quad \beta c=v, \quad \frac{\beta}{c}=\frac{v}{c^{2}}
$$

where we rewrote the first equations in terms of $\omega$ instead of $\omega / c$. The change in frequency due to motion is the basis of the Doppler effect. The invariance property (G.12) applied to the space-time and frequency-wavenumber four-vectors reads:

$$
\begin{equation*}
\omega^{\prime} t^{\prime}-\boldsymbol{k}^{\prime} \cdot \mathbf{r}^{\prime}=\omega t-\boldsymbol{k} \cdot \mathbf{r} \tag{G.15}
\end{equation*}
$$

This implies that a uniform plane wave remains a uniform plane wave in all reference frames moving at a constant velocity relative to each other. Similarly, the charge and current densities transform as follows:

$$
\begin{array}{|l}
\hline c \rho^{\prime}=\gamma\left(c \rho-\beta J_{z}\right)  \tag{G.16}\\
J_{z}^{\prime}=\gamma\left(J_{z}-\beta c \rho\right) \\
J_{x}^{\prime}=J_{x} \\
J_{y}^{\prime}=J_{y}
\end{array} \Leftrightarrow \begin{aligned}
& c \rho=\gamma\left(c \rho^{\prime}+\beta J_{z}^{\prime}\right) \\
& J_{z}=\gamma\left(J_{z}^{\prime}+\beta c \rho^{\prime}\right) \\
& J_{x}=J_{x}^{\prime} \\
& J_{y}=J_{y}^{\prime} \\
& \hline
\end{aligned}
$$

Because Eq. (G.5) implies that $L^{-T}=G L G$, we are led to define four-vectors that transform according to $L^{-T}$. Such four-vectors are referred to as being covariant. Given any contravariant 4 -vector $a$, we define its covariant version by $\bar{a}=G a$. This operation simply reverses the sign of the spatial part of $a$ :

$$
\bar{a}=G a=\left[\begin{array}{rr}
1 & 0  \tag{G.17}\\
0 & -I
\end{array}\right]\left[\begin{array}{c}
a_{0} \\
\boldsymbol{a}
\end{array}\right]=\left[\begin{array}{c}
a_{0} \\
-\boldsymbol{a}
\end{array}\right]
$$

The vector $\bar{a}$ transforms as follows:

$$
\begin{equation*}
\bar{a}^{\prime}=G a^{\prime}=G L a=(G L G)(G a)=L^{-T} \bar{a} \tag{G.18}
\end{equation*}
$$

where we used the property that $G^{2}=I_{4}$, the $4 \times 4$ identity matrix. The most important covariant vector is the four-dimensional gradient:

$$
\partial_{x}=\left[\begin{array}{l}
\partial_{\tau}  \tag{G.19}\\
\partial_{x} \\
\partial_{y} \\
\partial_{z}
\end{array}\right]=\left[\begin{array}{l}
\partial_{\tau} \\
\nabla
\end{array}\right]
$$

Because $x^{\prime}=L x$, it follows that $\partial_{x^{\prime}}=L^{-T} \partial_{x}$. Indeed, we have component-wise:

$$
\frac{\partial}{\partial x_{i}}=\sum_{j} \frac{\partial x_{j}^{\prime}}{\partial x_{i}} \frac{\partial}{\partial x_{j}^{\prime}}=\sum_{j} L_{j i} \frac{\partial}{\partial x_{j}^{\prime}} \quad \Rightarrow \quad \partial_{\mathrm{x}}=L^{T} \partial_{x^{\prime}} \quad \Rightarrow \quad \partial_{x^{\prime}}=L^{-T} \partial_{\mathrm{x}}
$$

For the $Z$-directed boost of Eq. (G.1), we have $L^{-T}=L^{-1}$, which gives:

$$
\begin{align*}
& \partial_{\tau^{\prime}}=\gamma\left(\partial_{\tau}+\beta \partial_{z}\right)  \tag{G.20}\\
& \partial_{z^{\prime}}=\gamma\left(\partial_{z}+\beta \partial_{\tau}\right) \\
& \partial_{x^{\prime}}=\partial_{x} \\
& \partial_{y^{\prime}}=\partial_{y}
\end{aligned} \Leftrightarrow \quad \begin{aligned}
& \partial_{\tau}=\gamma\left(\partial_{\tau^{\prime}}-\beta \partial_{z^{\prime}}\right) \\
& \partial_{z}=\gamma\left(\partial_{z^{\prime}}-\beta \partial_{\tau^{\prime}}\right) \\
& \partial_{x}=\partial_{x^{\prime}} \\
& \partial_{y}=\partial_{y^{\prime}}
\end{align*}
$$

The four-dimensional divergence of a four-vector is a Lorentz scalar. For example, denoting the current density four-vector by $J=\left[c \rho, J_{x}, J_{y}, J_{z}\right]^{T}$, the charge conservation law involves the four-dimensional divergence:

$$
\partial_{t} \rho+\nabla \cdot \boldsymbol{J}=\left[\partial_{\tau}, \partial_{x}, \partial_{y}, \partial_{z}\right]\left[\begin{array}{c}
c \rho  \tag{G.21}\\
J_{x} \\
J_{y} \\
J_{z}
\end{array}\right]=\partial_{x}^{T} J
$$

Under a Lorentz transformation, this remains invariant, and therefore, if it is zero in one frame it will remain zero in all frames. Using $\partial_{\mathrm{x}}^{T}=\partial_{\mathrm{x}^{\prime}}^{T} L$, we have:

$$
\begin{equation*}
\partial_{t} \rho+\nabla \cdot \boldsymbol{J}=\partial_{x}^{T} J=\partial_{x^{\prime}}^{T} L J=\partial_{x^{\prime}} J^{\prime}=\partial_{t^{\prime}} \rho^{\prime}+\nabla^{\prime} \cdot \boldsymbol{J}^{\prime} \tag{G.22}
\end{equation*}
$$

Although many quantities in electromagnetism transform like four-vectors, such as the space-time or the frequency-wavenumber vectors, the actual electromagnetic fields do not. Rather, they transform like six-vectors or rank-2 antisymmetric tensors.

A rank-2 tensor is represented by a $4 \times 4$ matrix, say $F$. Its Lorentz transformation properties are the same as the transformation of the product of a column and a row four-vector, that is, $F$ transforms like the quantity $a b^{T}$, where $a, b$ are column fourvectors. This product transforms like $a^{\prime} b^{\prime T}=L\left(a b^{T}\right) L^{T}$. Thus, a general second-rank tensor transforms as follows:

$$
\begin{equation*}
F^{\prime}=L F L^{T} \tag{G.23}
\end{equation*}
$$

An antisymmetric rank-2 tensor $F$ defines, and is completely defined by, two threedimensional vectors, say $\boldsymbol{a}=\left[a_{x}, a_{y}, a_{z}\right]^{T}$ and $\boldsymbol{b}=\left[b_{x}, b_{y}, b_{z}\right]^{T}$. Its matrix form is:

$$
F=\left[\begin{array}{cccc}
0 & -a_{x} & -a_{y} & -a_{z} \\
a_{x} & 0 & -b_{z} & b_{y} \\
a_{y} & b_{z} & 0 & -b_{x} \\
a_{z} & -b_{y} & b_{x} & 0
\end{array}\right]
$$

Given the tensor $F$, one may define its covariant version through $\bar{F}=G F G$, and its dual, denoted by $\tilde{F}$ and obtained by the replacements $\boldsymbol{a} \rightarrow \boldsymbol{b}$ and $\boldsymbol{b} \rightarrow-\boldsymbol{a}$, that is,

$$
\bar{F}=\left[\begin{array}{cccc}
0 & a_{x} & a_{y} & a_{z}  \tag{G.25}\\
-a_{x} & 0 & -b_{z} & b_{y} \\
-a_{y} & b_{z} & 0 & -b_{x} \\
-a_{z} & -b_{y} & b_{x} & 0
\end{array}\right], \quad \tilde{F}=\left[\begin{array}{cccc}
0 & -b_{x} & -b_{y} & -b_{z} \\
b_{x} & 0 & a_{z} & -a_{y} \\
b_{y} & -a_{z} & 0 & a_{x} \\
b_{z} & a_{y} & -a_{x} & 0
\end{array}\right]
$$

Thus, $\bar{F}$ corresponds to the pair $(-\boldsymbol{a}, \boldsymbol{b})$, and $\tilde{F}$ to $(\boldsymbol{b},-\boldsymbol{a})$. Their Lorentz transformation properties are:

$$
\begin{equation*}
\bar{F}^{\prime}=L^{-T} \bar{F} L^{-1}, \quad \tilde{F}^{\prime}=L \tilde{F} L^{T} \tag{G.26}
\end{equation*}
$$

Thus, the dual $\tilde{F}$ transforms like $F$ itself. For the $z$-directed boost of Eq. (G.1), it follows from (G.23) that the two vectors $\boldsymbol{a}, \boldsymbol{b}$ transform as follows:

$$
\begin{array}{ll}
a_{x}^{\prime}=\gamma\left(a_{x}-\beta b_{y}\right) & b_{x}^{\prime}=\gamma\left(b_{x}+\beta a_{y}\right) \\
a_{y}^{\prime}=\gamma\left(a_{y}+\beta b_{x}\right) & b_{y}^{\prime}=\gamma\left(b_{y}-\beta a_{x}\right)  \tag{G.27}\\
a_{z}^{\prime}=a_{z} & b_{z}^{\prime}=b_{z}
\end{array}
$$

These are obtained by equating the expressions:

$$
\begin{aligned}
& {\left[\begin{array}{cccc}
0 & -a_{x}^{\prime} & -a_{y}^{\prime} & -a_{z}^{\prime} \\
a_{x}^{\prime} & 0 & -b_{z}^{\prime} & b_{y}^{\prime} \\
a_{y}^{\prime} & b_{z}^{\prime} & 0 & -b_{x}^{\prime} \\
a_{z}^{\prime} & -b_{y}^{\prime} & b_{x}^{\prime} & 0
\end{array}\right]=} \\
= & {\left[\begin{array}{cccc}
\gamma & 0 & 0 & -\gamma \beta \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
-\gamma \beta & 0 & 0 & \gamma
\end{array}\right]\left[\begin{array}{cccc}
0 & -a_{x} & -a_{y} & -a_{z} \\
a_{x} & 0 & -b_{z} & b_{y} \\
a_{y} & b_{z} & 0 & -b_{x} \\
a_{z} & -b_{y} & b_{x} & 0
\end{array}\right]\left[\begin{array}{cccc}
\gamma & 0 & 0 & -\gamma \beta \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
-\gamma \beta & 0 & 0 & \gamma
\end{array}\right] }
\end{aligned}
$$

More generally, under the boost transformation (G.6), it can be verified that the components of $\boldsymbol{a}, \boldsymbol{b}$ parallel and perpendicular to $\boldsymbol{v}$ transform as follows:

$$
\begin{align*}
& \boldsymbol{a}_{\perp}^{\prime}=\gamma\left(\boldsymbol{a}_{\perp}+\boldsymbol{\beta} \times \boldsymbol{b}_{\perp}\right) \\
& \boldsymbol{b}_{\perp}^{\prime}=\gamma\left(\boldsymbol{b}_{\perp}-\boldsymbol{\beta} \times \boldsymbol{a}_{\perp}\right)  \tag{G.28}\\
& \boldsymbol{a}_{\|}^{\prime}=\boldsymbol{a}_{\|} \\
& \boldsymbol{b}_{\|}^{\prime}=\boldsymbol{b}_{\|} \\
&
\end{align*}
$$

$$
\gamma=\frac{1}{\sqrt{1-\beta^{2}}}, \quad \beta=|\boldsymbol{\beta}|, \quad \boldsymbol{\beta}=\frac{\boldsymbol{v}}{\boldsymbol{c}}
$$

Thus, in contrast to Eq. (G.11) for a four-vector, the parallel components remain unchanged while the transverse components change. A pair of three-dimensional vectors ( $\boldsymbol{a}, \boldsymbol{b}$ ) transforming like Eq. (G.28) is referred to as a six-vector.

It is evident also that Eqs. (G.28) remain invariant under the duality transformation $\boldsymbol{a} \rightarrow \boldsymbol{b}$ and $\boldsymbol{b} \rightarrow-\boldsymbol{a}$, which justifies Eq. (G.26). Some examples of (a,b) six-vector pairs defining an antisymmetric rank-2 tensor are as follows:

| $\boldsymbol{a}$ | $\boldsymbol{b}$ |
| :---: | ---: |
| $\boldsymbol{E}$ | $c \boldsymbol{B}$ |
| $c \boldsymbol{D}$ | $\boldsymbol{H}$ |
| $c \boldsymbol{P}$ | $-\boldsymbol{M}$ |

where $\boldsymbol{P}, \boldsymbol{M}$ are the polarization and magnetization densities defined through the relationships $\boldsymbol{D}=\epsilon_{o} \boldsymbol{E}+\boldsymbol{P}$ and $\boldsymbol{B}=\mu_{0}(\boldsymbol{H}+\boldsymbol{M})$. Thus, the $(\boldsymbol{E}, \boldsymbol{B})$ and $(\boldsymbol{D}, \boldsymbol{H})$ fields have the following Lorentz transformation properties:

$$
\begin{array}{|l}
\hline \boldsymbol{E}_{\perp}^{\prime}=\gamma\left(\boldsymbol{E}_{\perp}+c \boldsymbol{\beta} \times \boldsymbol{B}_{\perp}\right) \\
\boldsymbol{B}_{\perp}^{\prime}=\gamma\left(\boldsymbol{B}_{\perp}-\frac{1}{c} \boldsymbol{\beta} \times \boldsymbol{E}_{\perp}\right) \\
\boldsymbol{E}_{\|}^{\prime}=\boldsymbol{E}_{\|} \\
\boldsymbol{B}_{\|}^{\prime}=\boldsymbol{B}_{\|}
\end{array} \quad \begin{aligned}
& \boldsymbol{H}_{\perp}^{\prime}=\gamma\left(\boldsymbol{H}_{\perp}-c \boldsymbol{\beta} \times \boldsymbol{D}_{\perp}\right) \\
& \boldsymbol{D}_{\perp}^{\prime}=\gamma\left(\boldsymbol{D}_{\perp}+\frac{1}{c} \boldsymbol{\beta} \times \boldsymbol{H}_{\perp}\right)  \tag{G.30}\\
& \boldsymbol{H}_{\|}^{\prime}=\boldsymbol{H}_{\|} \\
& \boldsymbol{D}_{\|}^{\prime}=\boldsymbol{D}_{\|} \\
& \hline
\end{aligned}
$$

where we may replace $c \boldsymbol{\beta}=\boldsymbol{v}$ and $\boldsymbol{\beta} / \boldsymbol{c}=\boldsymbol{v} / c^{2}$. Note that the two groups of equations transform into each other under the usual duality transformations: $\boldsymbol{E} \rightarrow \boldsymbol{H}, \boldsymbol{H} \rightarrow-\boldsymbol{E}$, $\boldsymbol{D} \rightarrow \boldsymbol{B}, \boldsymbol{B} \rightarrow-\boldsymbol{D}$. For the $z$-directed boost of Eq. (G.1), we have from Eq. (G.30):

$$
\begin{align*}
& E_{x}^{\prime}=\gamma\left(E_{x}-c \beta B_{y}\right)  \tag{G.31}\\
& E_{y}^{\prime}=\gamma\left(E_{y}+c \beta B_{x}\right) \\
& B_{x}^{\prime}=\gamma\left(B_{x}+\frac{1}{c} \beta E_{y}\right) \\
& B_{y}^{\prime}=\gamma\left(B_{y}-\frac{1}{c} \beta E_{x}\right) \\
& E_{z}^{\prime}=E_{z} \\
& B_{z}^{\prime}=B_{z}
\end{aligned} \quad \begin{aligned}
& H_{x}^{\prime}=\gamma\left(H_{x}+c \beta D_{y}\right) \\
& H_{y}^{\prime}=\gamma\left(H_{y}-c \beta D_{x}\right) \\
& D_{x}^{\prime}=\gamma\left(D_{x}-\frac{1}{c} \beta H_{y}\right) \\
& D_{y}^{\prime}=\gamma\left(D_{y}+\frac{1}{c} \beta H_{x}\right) \\
& H_{z}^{\prime}=H_{z} \\
& D_{z}^{\prime}=D_{z} \\
& \hline
\end{align*}
$$

Associated with a six-vector $(\boldsymbol{a}, \boldsymbol{b})$, there are two scalar invariants: the quantities $(\boldsymbol{a} \cdot \boldsymbol{b})$ and $(\boldsymbol{a} \cdot \boldsymbol{a}-\boldsymbol{b} \cdot \boldsymbol{b})$. Their invariance follows from Eq. (G.28). Thus, the scalars $(\boldsymbol{E} \cdot \boldsymbol{B}),\left(\boldsymbol{E} \cdot \boldsymbol{E}-\boldsymbol{c}^{2} \boldsymbol{B} \cdot \boldsymbol{B}\right),(\boldsymbol{D} \cdot \boldsymbol{H}),\left(c^{2} \boldsymbol{D} \cdot \boldsymbol{D}-\boldsymbol{H} \cdot \boldsymbol{H}\right)$ remain invariant under Lorentz transformations. In addition, it follows from (G.30) that the quantity ( $\boldsymbol{E} \cdot \boldsymbol{D}-\boldsymbol{B} \cdot \boldsymbol{H}$ ) is invariant.

Given a six-vector ( $\boldsymbol{a}, \boldsymbol{b}$ ) and its dual ( $\boldsymbol{b},-\boldsymbol{a}$ ), we may define the following fourdimensional "current" vectors that are dual to each other:

$$
J=\left[\begin{array}{c}
\nabla \cdot \boldsymbol{a}  \tag{G.32}\\
\nabla \times \boldsymbol{b}-\partial_{\tau} \boldsymbol{a}
\end{array}\right], \quad \tilde{J}=\left[\begin{array}{c}
\nabla \cdot \boldsymbol{b} \\
-\nabla \times \boldsymbol{a}-\partial_{\tau} \boldsymbol{b}
\end{array}\right]
$$

It can be shown that both $J$ and $\tilde{J}$ transform as four-vectors under Lorentz transformations, that is, $J^{\prime}=L J$ and $\tilde{J}^{\prime}=L \tilde{J}$, where $J^{\prime}, \tilde{J}^{\prime}$ are defined with respect to the coordinates of the $S^{\prime}$ frame:

$$
J^{\prime}=\left[\begin{array}{c}
\nabla^{\prime} \cdot \boldsymbol{a}^{\prime}  \tag{G.33}\\
\nabla^{\prime} \times \boldsymbol{b}^{\prime}-\partial_{\tau^{\prime}} \boldsymbol{a}^{\prime}
\end{array}\right], \quad \tilde{J}^{\prime}=\left[\begin{array}{c}
\boldsymbol{\nabla}^{\prime} \cdot \boldsymbol{b}^{\prime} \\
-\nabla^{\prime} \times \boldsymbol{a}^{\prime}-\partial_{\tau^{\prime}} \boldsymbol{b}^{\prime}
\end{array}\right]
$$

The calculation is straightforward but tedious. For example, for the $z$-directed boost (G.1), we may use Eqs. (G.20) and (G.27) and the identity $\gamma^{2}\left(1-\beta^{2}\right)=1$ to show:

$$
\begin{aligned}
J_{x}^{\prime} & =\left(\nabla^{\prime} \times \boldsymbol{b}^{\prime}-\partial_{\tau^{\prime}} \boldsymbol{a}^{\prime}\right)_{x}=\partial_{y^{\prime}} b_{z}^{\prime}-\partial_{z^{\prime}} b_{y}^{\prime}-\partial_{\tau^{\prime}} a_{x}^{\prime} \\
& =\partial_{y} b_{z}-\gamma^{2}\left(\partial_{z}+\beta \partial_{\tau}\right)\left(b_{y}-\beta a_{x}\right)-\gamma^{2}\left(\partial_{\tau}+\beta \partial_{z}\right)\left(a_{x}-\beta b_{y}\right) \\
& =\partial_{y} b_{z}-\partial_{z} b_{y}-\partial_{\tau} a_{x}=\left(\boldsymbol{\nabla} \times \boldsymbol{b}-\partial_{\tau} \boldsymbol{a}\right)_{x}=J_{x}
\end{aligned}
$$

Similarly, we have:

$$
\begin{aligned}
J_{0}^{\prime} & =\nabla^{\prime} \cdot \boldsymbol{a}^{\prime}=\partial_{x^{\prime}} a_{x}^{\prime}+\partial_{y^{\prime}} a_{y}^{\prime}+\partial_{z^{\prime}} a_{z}^{\prime} \\
& =\gamma \partial_{x}\left(a_{x}-\beta b_{y}\right)+\gamma \partial_{y}\left(a_{y}+\beta b_{x}\right)+\gamma\left(\partial_{z}+\beta \partial_{\tau}\right) a_{z} \\
& =\gamma\left[\left(\partial_{x} a_{x}+\partial_{y} a_{y}+\partial_{z} a_{z}\right)-\beta\left(\partial_{x} b_{y}-\partial_{y} b_{x}-\partial_{\tau} a_{z}\right)\right]=\gamma\left(J_{0}-\beta J_{z}\right)
\end{aligned}
$$

In this fashion, one can show that $J$ and $\tilde{J}$ satisfy the Lorentz transformation equations (G.10) for a four-vector. To see the significance of this result, we rewrite Maxwell's equations, with magnetic charge and current densities $\rho_{m}, \boldsymbol{J}_{m}$ included, in the fourdimensional forms:

$$
\left[\begin{array}{c}
\nabla \cdot c \boldsymbol{D}  \tag{G.34}\\
\nabla \times \boldsymbol{H}-\partial_{\tau} c \boldsymbol{D}
\end{array}\right]=\left[\begin{array}{c}
c \rho \\
\boldsymbol{J}
\end{array}\right], \quad\left[\begin{array}{c}
\nabla \cdot c \boldsymbol{B} \\
-\nabla \times \boldsymbol{E}-\partial_{\tau} c \boldsymbol{B}
\end{array}\right]=\left[\begin{array}{c}
c \rho_{m} \\
\boldsymbol{J}_{m}
\end{array}\right]
$$

Thus, applying the above result to the six-vector ( $c \boldsymbol{D}, \boldsymbol{H}$ ) and to the dual of ( $\boldsymbol{E}, \boldsymbol{c} \boldsymbol{B}$ ) and assuming that the electric and magnetic current densities transform like fourvectors, it follows that Maxwell's equations remain invariant under Lorentz transformations, that is, they retain their form in the moving system:

$$
\left[\begin{array}{c}
\boldsymbol{\nabla}^{\prime} \cdot c \boldsymbol{D}^{\prime}  \tag{G.35}\\
\nabla^{\prime} \times \boldsymbol{H}^{\prime}-\partial_{\tau^{\prime} c} \boldsymbol{C} \boldsymbol{D}^{\prime}
\end{array}\right]=\left[\begin{array}{c}
c \rho^{\prime} \\
\boldsymbol{J}^{\prime}
\end{array}\right], \quad\left[\begin{array}{c}
\boldsymbol{\nabla}^{\prime} \cdot c \boldsymbol{B}^{\prime} \\
-\nabla^{\prime} \times \boldsymbol{E}^{\prime}-\partial_{\tau^{\prime}} c \boldsymbol{B}^{\prime}
\end{array}\right]=\left[\begin{array}{c}
c \rho_{m}^{\prime} \\
\boldsymbol{J}_{m}^{\prime}
\end{array}\right]
$$

The Lorentz transformation properties of the electromagnetic fields allow one to solve problems involving moving media, such as the Doppler effect, reflection and transmission from moving boundaries, and so on. The main technique for solving such problems is to transform to the frame (here, $S^{\prime}$ ) in which the boundary is at rest, solve the reflection problem in that frame, and transform the results back to the laboratory frame by using the inverse of Eq. (G.30).

This procedure was discussed by Einstein in his 1905 paper on special relativity in connection to the Doppler effect from a moving mirror. To quote [123]: "All problems in the optics of moving bodies can be solved by the method here employed. What is essential is that the electric and magnetic force of the light which is influenced by a moving body, be transformed into a system of co-ordinates at rest relatively to the body. By this means all problems in the optics of moving bodies will be reduced to a series of problems in the optics of stationary bodies."

## H. MATLAB Functions

The MATLAB functions are grouped by category. They are available from the web page: www.ece.rutgers.edu/~orfanidi/ewa.

## Multilayer Dielectric Structures

brewster - calculates Brewster and critical angles
fresne1 - Fresnel reflection coefficients for isotropic or birefringent media
n2r - refractive indices to reflection coefficients of M-layer structure
r2n - reflection coefficients to refractive indices of M-layer structure
multidiel - reflection response of a multilayer dielectric structure
omniband - bandwidth of omnidirectional mirrors and Brewster polarizers omniband2 - bandwidth of birefringent multilayer mirrors
snell - calculates refraction angles from Snell's law for birefringent media

## Quarter-Wavelength Transformers

| bkwrec - order-decreasing backward layer recursion - from a,b to $r$ |  |
| :--- | :--- |
| frwrec | - order-increasing forward layer recursion - from $r$ to A,B |
| chebtr | - Chebyshev broadband reflection1ess quarter-wave transformer |
| chebtr2 | - Chebyshev broadband reflectionless quarter-wave transformer |

chebtr2 - Chebyshev broadband reflectionless quarter-wave transformer
chebtr3 - Chebyshev broadband reflectionless quarter-wave transformer

## Dielectric Waveguides

| dguide | - TE modes in dielectric slab waveguide |
| :--- | :--- |
| dslab | - solves for the TE-mode cutoff wavenumbers in a dielectric slab |

## Transmission Lines

| g2z | - reflection coefficient to impedance transformation |
| :--- | :--- |
| z2g | - impedance to reflection coefficient transformation |
| lmin | - find locations of voltage minima and maxima |
| mstripa | - microstrip analysis (calculates Z, eff from w/h) |
| mstripr | - microstrip synthesis with refinement (calculates w/h from Z) |
| mstrips | - microstrip synthesis (calculates w/h from Z) |
| multiline - reflection response of multi-segment transmission line |  |
| swr | - standing wave ratio |
| tsection | - T-section equivalent of a length-1 transmission line segment |
| gprop | - reflection coefficient propagation |
| vprop | - wave impedance propagation |
| zprop | - wave impedance propagation |

## Impedance Matching

qwt1 - quarter wavelength transformer with series segment
qwt2 - quarter wavelength transformer with $1 / 8$-wavelength shunt stub qwt3 - quarter wavelength transformer with shunt stub of adjustable length
dualband - two-section dual-band Chebyshev impedance transformer
dualbw - two-section dual-band transformer bandwidths
stub1 - single-stub matching
stub2 - double-stub matching
stub3 - triple-stub matching
onesect
pi2t - Pi to T transformation
t2pi - Pi to $T$ transformation
7match - L-section reactive conjugate matching network
pmatch - Pi-section reactive conjugate matching network

## S-Parameters

gin - input reflection coefficient in terms of S-parameters
gout - output reflection coefficient in terms of S-parameters
nfcirc - constant noise figure circle
nfig - noise figure of two-port
sgain - transducer, available, and operating power gains of two-port
sgcirc - stability and gain circles
smat - S-parameters to S-matrix
smatch - simultaneous conjugate match of a two-port
smith - draw basic Smith chart
smithcir - add stability and constant gain circles on Smith chart
sparam - stability parameters of two-port
circint - circle intersection on Gamma-plane
circtan - point of tangency between the two circles

## Linear Antenna Functions

| dipole | - gain of center-fed linear dipole of length $L$ |
| :--- | :--- |
| travel | - gain of traveling-wave antenna of length L |
| vee | - gain of traveling-wave vee antenna |
| rhombic | - gain of traveling-wave rhombic antenna |
| dmax | - computes directivity and beam solid angle of $g($ th $)$ gain |
| hallen | - solve Hallen's integral equation with delta-gap input |
| hallen2 | - solve Hallen's integral equation with arbitrary incident E-field |
| hallen3 | - solve Hallen's equation for 2D array of identical linear antennas |
| hallen4 | - solve Hallen's equation for 2D array of non-identical linear antennas |
| pockling | - solve Pocklington's integral equation for linear antenna |
| king | - King's 3-term sinusoidal approximation |
| kingeval | - evaluate King's 3-term sinusoidal current approximation |
| kingfit | - fits a sampled current to King's 2-term sinusoidal approximation |
| gain2 | - normalized gain of arbitrary 2D array of linear sinusoidal antennas |
| gain2h | - gain of 2D array of linear antennas with Hallen currents |
| imped | - mutual impedance between two parallel standing-wave dipoles |
| impedmat | - mutual impedance matrix of array of parallel dipole antennas |
| yagi | - simplified Yagi-Uda array design |

yagi - simplified Yagi-Uda array design

## Aperture Antenna Functions

| diffint diffr dsinc | - generalized Fresne1 diffraction integral <br> - knife-edge diffraction coefficient <br> - the double-sinc function $\cos (p i * x) /(1-4 * x \wedge 2)$ |
| :---: | :---: |
| fcs | - Fresnel integrals $C(x)$ and $S(x)$ |
| fcs2 | - type-2 Fresne1 integrals C2(x) and S2(x) |
| hband | horn antenna 3-dB width |
| heff | - aperture efficiency of horn antenna |
| hgain | - horn antenna H-plane and E-plane gains |
| hopt | - optimum horn antenna design |
| hsigma | - optimum sigma parametes for horn antenn |

## Antenna Array Functions

array - gain computation for 1D equally-spaced isotropic array bwidth - beamwidth mapping from psi-space to phi-space
binomial - binomial array weights
dolph - Dolph-Chebyshev array weights
dolph2 - Riblet-Pritchard version of Dolph-Chebyshev
dolph3 - DuHamel version of endfire Dolph-Chebyshev
multibeam - multibeam array design
scan - scan array with given scanning phase
sector - sector beam array design
steer - steer array towards given angle
taylor - Taylor-Kaiser window array weights
uniform - uniform array weights
woodward - Woodward-Lawson-Butler beams
chebarray - Bresler's Chebyshev array design method (written by P. Simon)

## Miscellaneous Utility Functions



## Gain Plotting Functions

| abp | - polar gain plot in absolute units |
| :--- | :--- |
| abz | - azimuthal gain plot in absolute units |
| ab2p | - polar gain plot in absolute units - 2*pi angle range |
| abz2 | - azimuthal gain plot in absolute units - 2pi angle range |
|  |  |
| dbp | - polar gain plot in dB |
| dbz | - azimuthal gain plot in dB |
| dbp2 | - polar gain plot in $\mathrm{dB}-2 *$ pi angle range |
| dbz2 | - azimuthal gain plot in dB - 2pi angle range |
|  |  |
| abadd | - add gain in absolute units |
| abadd2 | - add gain in absolute units - 2pi angle range |
| dbadd | - add gain in dB |
| dbadd2 | - add gain in dB - 2pi angle range |
| addbwp | - add 3-dB angle beamwidth in polar plots |
| addbwz | - add 3-dB angle beamwidth in azimuthal plots |
| addcirc | - add grid circle in polar or azimuthal plots |
| addline | - add grid ray line in azimuthal or polar plots |
| addray | - add ray in azimuthal or polar plots |


[^0]:    ${ }^{\dagger}$ International Telecommunication Union.

[^1]:    ${ }^{\dagger}$ Commission Internationale de l’Eclairage (International Commission on Illumination.)
    ${ }^{\ddagger} 1 \mathrm{THz}=10^{12} \mathrm{~Hz}, 1 \mathrm{PHz}=10^{15} \mathrm{~Hz}, 1 \mathrm{EHz}=10^{18} \mathrm{~Hz}$.

